

# BCS superconductivity in metallic nanograins: Finite-size corrections, low energy excitations, and robustness of shell effects

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We combine the BCS self-consistency condition, a semiclassical expansion for the spectral density and interaction matrix elements to describe analytically how the superconducting gap depends on the size and shape of a 2d and 3d superconducting grain. In chaotic grains mesoscopic fluctuations of the matrix elements lead to a smooth dependence of the order parameter on the excitation energy. In the integrable case we find shell effects i. e. for certain values of the electron number  $N$  a small change in  $N$  leads to large changes in the energy gap. With regard to possible experimental tests we provide a detailed analysis of the dependence of the gap on the coherence length and the robustness of shell effects under small geometrical deformations.

Finite size effects are well documented [1] in fermionic interacting systems such as atomic nuclei and atomic clusters. It is also well established [2, 3] that the more symmetric the system is, the stronger are these corrections. For instance, the existence of magic numbers signaling the presence of a particularly stable nucleus has its origin in the gap between the ground state and the first excited states caused by the high degree of symmetry of the system.

In the field of mesoscopic superconductivity, the study of finite size effects also has a long history. Already fifty years ago, Anderson noted [4] that superconductivity should break down in small metallic grains when the single particle level spacing at the Fermi energy is comparable to the bulk superconducting gap. In the sixties the size dependence of the critical temperature and the superconducting gap were studied in for a rectangular grain in [5] and for a nanoslab in [6]. Thermodynamical properties of superconducting grains were investigated in [7]. Results of these papers are restricted to rectangular grains, and superconductivity is described by the Bardeen, Cooper, and Schrieffer (BCS) theory [8].

The experiments by Ralph, Black, and Tinkham in the mid nineties [9] on Al nanograins of typical size  $L \sim 3 - 13$  nm showed that the excitation gap is sensitive to even-odd effects. More recently it has been observed [10] that the critical temperature of superconducting ultrathin lead films oscillates when the film thickness is slightly increased. These results have further stimulated the interest in ultrasmall superconductors [11–16]. For instance, pairing, not necessarily BCS, in a harmonic oscillator potential was investigated in [13]. The critical temperature and the superconducting gap for a nanowire were reported in [14] by solving numerically the Bogoli-

ubov - de Gennes equations. In [15] the superconducting gap and low energy excitation energies in a rectangular grain were computed numerically within the Richardson model [16]. Shell effects in superconducting grains with radial symmetry were studied in [17, 18]. Moreover, recent experiments on Al grains were interpreted [19] as evidence that shell effects can drive critical temperatures in these grains above 100 K. Mesoscopic corrections to the BCS energy gap were also considered in [20, 21].

We note that if the mean single particle level spacing is larger than the bulk superconducting gap, the BCS formalism breaks down. However, an analytical treatment is still possible [22] with the help of an exactly solvable model introduced by Richardson [16] in the context of nuclear physics. In particular, finite-size corrections to the predictions of the BCS theory have been recently studied in [23–27].

Despite this progress, a theory that accounts for all relevant mesoscopic effects in superconducting grains has not emerged so far. The Richardson model alone cannot provide the foundation for such a theory as it does not allow for mesoscopic spatial fluctuations of the single particle states. In the present paper, for the particular cases of chaotic and rectangular shaped grains, we develop such a theory based on the BCS theory and semiclassical techniques. This formalism permits a systematic analytical evaluation of the low energy spectral properties of superconducting nanograins in terms of their size and shape. Leading finite size corrections to the BCS mean field can also be taken into account in our approach, see [28] for further details. Results for 3d grains were also previously published in [28]. Here we discuss both the 2d and 3d cases as well as provide a more detailed account of the techniques utilized. Moreover, we study the dependence

of the mesoscopic BCS order parameter (superconducting gap) on the coherence length, and the robustness of shell effects.

For chaotic grains, we show that the order parameter is a universal function of the single particle energy, i.e. it is independent of the particular details of the grain. The mesoscopic fluctuations of the matrix elements of the two-body interactions between single particle eigenstates are responsible for most of the deviations from the bulk limit. For integrable grains, we find that the superconducting gap is strongly sensitive to shell effects. Namely, a small modification of the grain size or number of electrons inside can substantially affect its value. Throughout the paper we study clean (ballistic) grains. The mean field potential is thus an infinite well of the form of the grain. We restrict ourselves to system sizes such that the mean level spacing around the Fermi energy is smaller than the bulk gap, so that the BCS formalism is still a good approximation. For the superconducting Al grains studied by Tinkham and coworkers [9], this corresponds to sizes  $L > 5$  nm.

Our results are therefore valid in the region,  $k_F L \gg 1$  (limit of validity of the semiclassical approximation [4, 23]),  $\delta/\Delta_0 < 1$  (limit of validity of the BCS theory), and  $l \gg \xi \gg L$  (condition of quantum coherence). Here  $k_F$ ,  $\xi = \hbar v_F/\Delta_0$ ,  $l$ ,  $\delta$ ,  $\Delta_0$  are the Fermi wave vector, the superconducting coherence length, the coherence length of the single particle problem, the average single particle level spacing, and the bulk gap. The Fermi velocity is  $v_F = \hbar k_F/m$ . Conditions  $k_F L \gg 1$  and  $\delta/\Delta_0 < 1$  hold for Al grains of size  $L \gtrsim 5$  nm. Further, in Al grains  $\xi \approx 1600$  nm and  $l > 10^4$  nm at temperatures  $T \leq 4$  K [24]. Therefore, the above region is well accessible to experiments.

## I. THE SUPERCONDUCTING GAP IN THE BCS THEORY

Throughout the paper pairing between electrons is described by the BCS Hamiltonian,

$$H = \sum_{n\sigma} \epsilon_n c_{n\sigma}^\dagger c_{n\sigma} - \sum_{n,n'} I_{n,n'} c_{n\uparrow}^\dagger c_{n\downarrow}^\dagger c_{n'\downarrow} c_{n'\uparrow},$$

where  $c_{n\sigma}$  annihilates an electron of spin  $\sigma$  in state  $n$ ,

$$I_{n,n'} \equiv I(\epsilon_n, \epsilon_{n'}) = \lambda V \delta \int \psi_n^2(\vec{r}) \psi_{n'}^2(\vec{r}) d\vec{r} \quad (1)$$

are matrix elements of a short-range electron-electron interaction,  $\lambda$  is the BCS coupling constant.  $\psi_n$  and  $\epsilon_n$  are the eigenstates and eigenvalues of a free particle of effective mass  $m$  in a clean grain of volume (area)  $V$  ( $A$ ). Eigenvalues  $\epsilon_n$  are measured from the Fermi energy  $\epsilon_F$ . In this notation the mean level spacing is  $\delta = 1/\nu_{\text{TF}}(0)$ , where  $\nu_{\text{TF}}(0)$  is the spectral density at the Fermi energy in the Thomas-Fermi approximation.

The BCS order parameter is defined as

$$\Delta_n \equiv \Delta(\epsilon_n) = \sum_n I_{n,n'} \langle c_{n'\uparrow}^\dagger c_{n'\downarrow}^\dagger \rangle.$$

Within BCS theory, it is determined by the following self-consistency equation[25]:

$$\Delta_n = \frac{1}{2} \sum_{|\epsilon_{n'}| < \epsilon_D} \frac{\Delta_{n'} I_{n,n'}}{\sqrt{\epsilon_{n'}^2 + \Delta_{n'}^2}}, \quad (2)$$

where  $\epsilon_D$  is the Debye energy. This result is obtained in the grand canonical approximation [8]. Note that, the BCS order parameter  $\Delta_n$  is an explicit function of the single-particle energy  $\epsilon_n$  since the matrix elements  $I(\epsilon, \epsilon')$  are energy dependent.

Introducing the exact density of single-particle states  $\nu(\epsilon') = \sum_{n'} \delta(\epsilon' - \epsilon_{n'})$ , one can write Eq. (2) in integral form,

$$\Delta(\epsilon) = \frac{1}{2} \int_{-\epsilon_D}^{\epsilon_D} \frac{\Delta(\epsilon') I(\epsilon, \epsilon')}{\sqrt{\epsilon'^2 + \Delta^2(\epsilon')}} \nu(\epsilon') d\epsilon'. \quad (3)$$

The gap equation (3) will be the main subject of our interest. As soon as the order parameter  $\Delta(\epsilon)$  is known, the low lying (single-particle) excitation spectrum,  $E = \sqrt{\Delta(\epsilon)^2 + \epsilon^2}$ , is also determined.

In the large volume (area) limit, the spectral density, to leading order, is given by the Thomas-Fermi expression

$$\nu_{\text{TF}}(\epsilon') = 2 \times \begin{cases} \frac{V}{4\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \sqrt{\epsilon' + \epsilon_F}, & \text{for 3d} \\ \frac{A}{4\pi} \left(\frac{2m}{\hbar^2}\right), & \text{for 2d,} \end{cases} \quad (4)$$

where the factor two in front stands for spin degeneracy. In addition, in the bulk limit the matrix elements (1) for chaotic grains are simply  $I(\epsilon, \epsilon') = \lambda \delta$  as a consequence of quantum ergodicity. The gap is then energy independent  $\Delta(\epsilon) = \Delta_0$ , and Eq. (2) yields the BCS bulk result,

$$\Delta_0 = 2\epsilon_D e^{-\frac{1}{\lambda}}. \quad (5)$$

As the volume of the grain decreases, both  $\nu(\epsilon')$  and  $I(\epsilon, \epsilon')$  deviate from the bulk limit. In this region a more general approach to solve Eq. (3) is needed.

Since we are interested in the regime of many particles ( $\nu_{\text{TF}}(0)\epsilon_F \gg 1$ ), an appropriate tool is the semiclassical approximation in general and periodic orbit theory [29] in particular (see the Appendix for an introduction). These techniques yield closed expressions for  $\nu(\epsilon')$  and  $I(\epsilon, \epsilon')$  in terms of quantities from the classical dynamics of the system, which allows us to calculate analytically the resulting superconducting gap. Such explicit expressions for the superconducting gap enable us to study deviations from the BCS theory, the spatial dependence of the gap, and the relevance of shell effects in realistic, not perfectly symmetric grains.

Our general strategy can be summarized as follows:

1. Use semiclassical techniques to compute the spectral density  $\nu(\epsilon') = \sum_{n'} \delta(\epsilon' - \epsilon_{n'})$  and  $I(\epsilon, \epsilon')$  as series in the small parameter  $1/k_F L$ , where  $k_F$  is the Fermi wave-vector and  $L \simeq V^{1/3} (\simeq A^{1/2})$  is the linear size of the grain (section II and Appendix).
2. Solve the BCS gap equation (2) order by order in  $1/k_F L$  (Section III).
3. Study the impact of small deformations of the shape of a symmetric grain on the gap in realistic models of the grain (Section IV).

## II. SEMICLASSICAL APPROXIMATION FOR THE DENSITY OF STATES AND INTERACTION MATRIX ELEMENTS.

The first step to solve the gap equation is to find explicit expressions for the spectral density  $\nu(\epsilon')$  and the interaction matrix elements  $I(\epsilon, \epsilon')$  as series in a small parameter  $1/k_F L$ . While the semiclassical approximation for the spectral density has been known for a long time [29], the calculation for the matrix elements has only recently attracted some attention [28, 30]. Here we state the results and refer the reader to the Appendix for details.

In the semiclassical approximation (see Appendix 1), the spectral density is given by

$$\nu(\epsilon') \simeq \nu_{\text{TF}}(0) [1 + \bar{g}(0) + \tilde{g}_l(\epsilon')], \quad (6)$$

with a monotonous  $\bar{g}(\epsilon')$  and oscillatory  $\tilde{g}(\epsilon')$  (as functions of system size) parts. The notation  $\bar{g}(\epsilon = 0)$  means that  $\bar{g}$  is evaluated at the Fermi energy. This contribution is given by the Weyl expansion [1],

$$\bar{g}(0) = \begin{cases} \pm \frac{S\pi}{4k_F V} + \frac{2\mathcal{C}}{k_F^2 V}, & \text{3d,} \\ \pm \frac{\mathcal{L}}{2k_F A}, & \text{2d,} \end{cases} \quad (7)$$

for Dirichlet (−) or Neumann (+) boundary conditions. In Eq. (7),  $S$  is the surface area of the 3d cavity and  $\mathcal{C}$  its mean curvature, while  $\mathcal{L}$  is the perimeter in the 2d case.

The oscillatory contribution to the density of states is given to leading order by the Gutzwiller trace formula [29],

$$\tilde{g}_l(\epsilon') = \Re \left\{ \begin{aligned} & \frac{2\pi}{k_F^2 V} \sum_p^l A_p e^{i[k_F L_p + \beta_p]} e^{i \frac{\epsilon'}{2\epsilon_F} k_F L_p} & \text{3d,} \\ & \frac{2}{k_F A} \sum_p^l A_p e^{i[k_F L_p + \beta_p]} e^{i \frac{\epsilon'}{2\epsilon_F} k_F L_p} & \text{2d.} \end{aligned} \right. \quad (8)$$

The summation over classical periodic orbits ( $p$ ) with length  $L_p$  only includes orbits shorter than the quantum coherence length  $l$  of the single-particle problem. The semiclassical amplitude  $A_p$  and phase  $\beta_p$  in Eq. (8) can also be computed explicitly using the knowledge of periodic orbits.

The calculation of the interaction matrix elements  $I(\epsilon, \epsilon')$  is more complicated as it requires information

about classical dynamics beyond periodic orbits. For a chaotic cavity the final result (Appendix 2),

$$I(\epsilon, \epsilon') = \begin{cases} \frac{\lambda}{V} \left[ 1 + \bar{I}_{3d}^{\text{short}}(0) - \frac{\pi^2 S^2}{16k_F^2 V^2} + \bar{I}_{\text{dg}}^{\text{long}}(0, \epsilon - \epsilon') \right] & \text{3d,} \\ \frac{\lambda}{A} \left[ 1 + \bar{I}_{2d}^{\text{short}}(0, \epsilon - \epsilon') + \bar{I}_{\text{dg}}^{\text{long}}(0, \epsilon - \epsilon') \right] & \text{2d,} \end{cases} \quad (9)$$

has two types of contributions. Identical pairs of short classical trajectories hitting the boundary once give

$$\begin{cases} \bar{I}_{3d}^{\text{short}}(0) = \frac{\pi S}{4k_F V} & \text{3d,} \\ \bar{I}_{2d}^{\text{short}}(0, \epsilon - \epsilon') = \frac{\mathcal{L}}{k_F A} \left[ C' + \frac{\text{Si}(4k_F L)}{\pi} \right] \\ + \frac{\mathcal{L}}{2\pi k_F A} \left[ \text{Ci} \left( \frac{4(\epsilon - \epsilon') k_F L}{\epsilon_F} \right) - \text{Ci} \left( \frac{2(\epsilon - \epsilon')}{\epsilon_F} \right) \right] & \text{2d,} \end{cases} \quad (10)$$

with  $C' = 0.339\dots$  a numerical constant given in the Appendix, and  $\text{Ci}(x)$  the cosine-integral function.

In the so-called diagonal approximation (see Appendix 2) the contribution of longer classical trajectories is

$$\bar{I}_{\text{dg}}^{\text{long}}(\epsilon_F, \epsilon - \epsilon') = \begin{cases} \frac{1}{V} \Pi_l \left( \frac{\epsilon - \epsilon'}{\epsilon_F} \right) & \text{3d,} \\ \frac{1}{A} \Pi_l \left( \frac{\epsilon - \epsilon'}{\epsilon_F} \right) & \text{2d,} \end{cases} \quad (11)$$

where

$$\Pi_l(w) = \int \sum_{\gamma}^l D_{\gamma}^2 \cos[wk_F L_{\gamma}(\vec{r})] d\vec{r} \quad (12)$$

is an integrated sum over trajectories  $\gamma(\vec{r})$  starting and ending at position  $\vec{r}$ . As detailed in Appendix 2, due to the ergodicity of the chaotic classical systems, in the limit  $l \gg L$ , Eq. (12) simplifies to

$$\Pi_{l \gg L}(w) = \begin{cases} \frac{4\pi^2 \sin(wk_F l)}{k_F^2 w} & \text{3d,} \\ \frac{4}{k_F^2} \frac{\sin(wk_F l)}{w} & \text{2d.} \end{cases} \quad (13)$$

For integrable grains there is no universal expression for  $I(\epsilon, \epsilon')$ . We restrict ourselves to the rectangular geometry where the matrix elements are energy independent.

Using the knowledge of  $\nu(\epsilon')$  and  $I(\epsilon, \epsilon')$  as series in  $1/k_F L$ , we solve the gap equation (3) in different situations of interest. The resulting gap function, in general, depends the single-particle energy  $\epsilon$ , the size of the system, and the number of particles (or, equivalently, Fermi energy  $\epsilon_F$ ).

## III. SOLUTION OF THE GAP EQUATION IN THE SEMICLASSICAL REGIME

In this section we solve the gap equation Eq. (3) for  $\Delta(\epsilon)$ . For a rectangular box in two and three dimensions the gap equation is algebraic, since  $\Delta(\epsilon) = \Delta$  is energy

independent. In the chaotic case, however, we get an integral equation due to the energy dependence of the interaction matrix elements. As we will see, both cases can be solved analytically order by order in  $1/k_FL$ .

### A. Rectangular box in two and three dimensions

For the rectangular box the matrix elements are  $I(\epsilon, \epsilon') = \prod_{i=x,y,z} (1 + \delta_{\epsilon_i, \epsilon'_i}/2)/V$  where  $\epsilon_i = \hbar^2 k_i^2$  and  $p_i = \hbar k_i$  is the conserved momentum in the  $i = x, y, z$  direction. The enhancement of the matrix elements for  $\epsilon_i = \epsilon'_i$  has a negligible effect in the gap equation since the number of levels in the interacting region is always  $2E_D/\delta \gg 1$ . As a consequence  $I(\epsilon, \epsilon') \approx 1/V$ , the gap does not depend on the energy,  $\Delta(\epsilon) = \Delta$ , and satisfies the equation

$$\frac{2}{\lambda} = \int_{-\epsilon_D}^{\epsilon_D} \frac{1 + \bar{g}(0) + \tilde{g}_l(\epsilon')}{\sqrt{\epsilon'^2 + \Delta^2}} d\epsilon', \quad (14)$$

where  $\bar{g}(0)$  for a 3d rectangular box is given by Eq. (7) without the curvature term.

Using Eq. (7) for  $\bar{g}(0)$  and Eq. (8) for  $\tilde{g}_l(\epsilon')$  (from now on we drop the subscript  $l$  to simplify the notation), and taking into account the scaling of each contribution with  $1/k_FL$  as described in the Appendix, we look for a solution of the gap equation (14) for the 3d case in the following form:

$$\Delta = \Delta_0(1 + f^{(1)} + f^{(3/2)} + f^{(2)}), \quad (15)$$

where  $f^{(n)} \propto 1/(k_FL)^n$ . Substituting  $\Delta$  into Eq. (14), expanding in powers of  $1/k_FL$ , and equating the coefficients at each power, we obtain an explicit expression for  $f^{(i)}$

$$\begin{aligned} \lambda f^{(1)} &= \left[ \bar{g}(0) + \frac{\lambda}{2} \int_{-\epsilon_D}^{\epsilon_D} \frac{\tilde{g}^{(3)}(\epsilon')}{\sqrt{\epsilon'^2 + \Delta_0^2}} d\epsilon' \right], \\ \lambda f^{(3/2)} &= \sum_{i,j \neq i}^3 \frac{\lambda}{2} \int_{-\epsilon_D}^{\epsilon_D} \frac{\tilde{g}_{i,j}^{(2)}(\epsilon')}{\sqrt{\epsilon'^2 + \Delta_0^2}} d\epsilon', \end{aligned} \quad (16)$$

$$\begin{aligned} \lambda f^{(2)} &= \sum_i^3 \frac{\lambda}{2} \int_{-\epsilon_D}^{\epsilon_D} \frac{\tilde{g}_i^{(1)}(\epsilon')}{\sqrt{\epsilon'^2 + \Delta_0^2}} d\epsilon' \\ &+ f^{(1)} \left( f^{(1)} - \bar{g}(0) \right) \\ &- f^{(1)} \sum_i \frac{\Delta_0^2}{2} \int_{-\epsilon_D}^{\epsilon_D} \frac{\tilde{g}_i^{(1)}(\epsilon')}{(\epsilon'^2 + \Delta_0^2)^{3/2}} d\epsilon', \end{aligned} \quad (17)$$

where  $\tilde{g}^{(k)} \propto (k_FL)^{-k}$  denotes the oscillating part of the spectral density. Explicit expressions for  $\tilde{g}^{(k)}$ ,  $\tilde{g}_i^{(k)}$ , and  $\tilde{g}_{i,j}^{(k)}$  for a rectangular box in terms of periodic orbits can be found in the Appendix and also in Ref. [1].

Equations (16) and (17) can be further simplified by the following argument. After we express  $\tilde{g}^{(3)}$ ,  $\tilde{g}^{(2)}$  and

$\tilde{g}^{(1)}$  in terms of a sum over periodic orbits, the integration over  $\epsilon'$  can be explicitly performed. The resulting expression is again an expansion in terms of periodic orbits with two peculiarities: a) the spectral density is evaluated at the Fermi energy and b) in the limit  $\epsilon_D \gg \Delta_0$  the contribution of an orbit of period  $L_p$  is weighted with the function

$$W(L_p/\xi) = \frac{\lambda}{2} \int_{-\infty}^{\infty} \frac{\cos(L_p t/\xi)}{\sqrt{1+t^2}} dt. \quad (18)$$

This cutoff function is characteristic of the BCS theory as opposed to the smoothing due to temperature or inelastic scattering (recall that in this paper we assume that the single-particle coherence length  $l$  is much larger than superconducting coherence length  $\xi$ ). In a similar fashion, the last term in  $f^{(2)}$  is weighted with

$$W_{3/2}(L_p/\xi) = \frac{\Delta_0^2}{2} \int_{-\infty}^{\infty} \frac{\cos(L_p t/\xi)}{(1+t^2)^{3/2}} dt.$$

The effect of  $W_{3/2}(L_p/\xi)$  is, again, to exponentially suppress the contribution of periodic orbits longer than  $\xi$ . Therefore the sum over periodic orbits in the definition of the spectral density is effectively restricted to orbits with lengths of the order or smaller than the superconducting coherence length  $\xi$ .

Following standard semiclassical approximations, we introduce  $\tilde{g}_\xi(0)$  as a spectral density evaluated at the Fermi energy with a cutoff function that suppresses the contribution of orbits of length  $L_p > \xi$ . With these definitions, we get

$$\begin{aligned} \lambda f^{(1)} &= \left[ \bar{g}(0) + \tilde{g}_\xi^{(3)}(0) \right], \\ \lambda f^{(3/2)} &= \sum_{i,j \neq i}^3 \tilde{g}_{i,j\xi}^{(2)}(0), \\ \lambda f^{(2)} &= \sum_i^3 \tilde{g}_{i\xi}^{(1)}(0) \\ &+ f^{(1)} \left[ f^{(1)} - \bar{g}(0) - \sum_i^3 \tilde{g}_{i\xi}^{(1)}(0) \right]. \end{aligned} \quad (19)$$

Eq. (19) is our final result for the finite size corrections to the gap function for a 3d rectangular box. As expected, it is expressed in terms of classical quantities such as the volume, surface, and periodic orbits of the grain.

In Fig. 1 we compare the analytical result (15) and (19) to the exact numerical solution of the gap equation for a superconducting Al grain of volume  $\approx 600 \text{ nm}^3$  with parameters  $\lambda \sim 0.18$ ,  $\Delta_0 = 0.24 \text{ meV}$  and  $\epsilon_F \sim 11 \text{ eV}$  similar to the ones utilized in the experiments of Ref. 9.

The following argument can shed light on our results. The density of states cannot be pulled out of the energy integration in the gap equation (14) unless it is smoothed. However, this is exactly what our result Eq. (19) means, since truncating the sums is equivalent to smoothing the energy dependence. We conclude that our result Eq. (14)

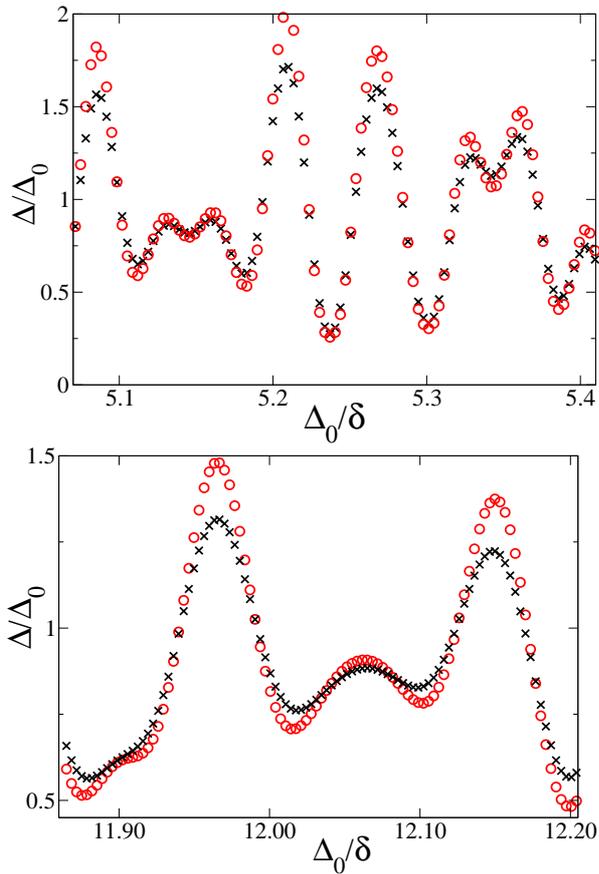


FIG. 1: Superconducting order parameter  $\Delta$  in units of the bulk gap  $\Delta_0$  for a cubic Al grain as a function of the ratio  $\Delta_0/\delta$ , where  $\delta$  is the mean level spacing. Crosses correspond to the exact numerical solution of the gap equation Eq.(2), while the circles represent the semiclassical analytical result Eqs. (15) and (19).

should be similar to the standard BCS solution in the bulk,  $\Delta_0 = 2\epsilon_D e^{-1/\lambda}$ , with the substitution  $\lambda \rightarrow \lambda(1 + \bar{g}(0) + \tilde{g}_\xi(0))$ . Indeed, an expansion of this expression in  $1/k_F L$  gives exactly Eq. (19).

In order to simplify notation from now on we will drop the subscript  $\xi$  in the spectral density  $\tilde{g}_\xi$  smoothed by the cutoff function  $W(L_p/\xi)$ . In 2d we find,

$$\Delta = \Delta_0(1 + f^{(1/2)} + f^{(1)}), \quad (20)$$

with

$$\begin{aligned} \lambda f^{(1/2)} &= \tilde{g}_{1,2}^{(2)}(0) \\ \lambda f^{(1)} &= \bar{g}(0) + \sum_{i=1,2} \tilde{g}_i^{(1)}(0) \\ &\quad + \frac{1-\lambda}{\lambda} [\tilde{g}_{1,2}^{(2)}(0)]^2. \end{aligned} \quad (21)$$

The sums implicit in  $\tilde{g}_i, \tilde{g}_{i,j}$  are smoothly truncated by the same weight function  $W(L_p/\xi)$ . Similar to the 3d case, the above result can also be obtained by expanding

the bulk expression for the gap with the full density of states in  $1/(k_F L)$ . We note that, contrary to the 3d case, in 2d grains, oscillatory contributions to the density of states are of leading order.

### B. 3d chaotic cavity

The energy dependence of the interaction matrix elements,  $I(\epsilon, \epsilon')$ , in this case is given by Eqs. (9–13), i.e.

$$I(\epsilon, \epsilon') = \frac{\lambda}{V} \left[ 1 + \frac{\pi S}{4k_F V} - \frac{\pi^2 S^2}{16k_F^2 V^2} + \frac{1}{V} \Pi_l \left( \frac{\epsilon - \epsilon'}{\epsilon_F} \right) \right],$$

where

$$\Pi_l(w) = \frac{4\pi^2}{k_F^3} \frac{\sin(k_F l w)}{w}. \quad (22)$$

The details of the calculation based on the semiclassical approximation for Green's functions can be found in Appendix 2.

The above expression for  $I(\epsilon, \epsilon')$  together with the semiclassical expression for the spectral density (8) are the starting point for the calculation of the superconducting order parameter. The energy dependence of the matrix elements implies a gap equation of integral type and, most importantly, that the order parameter itself depends on the energy. Based on the  $1/k_F L$  dependence of the different contributions to  $I(\epsilon, \epsilon')$ , we write

$$\Delta(\epsilon) = \Delta_0 \left[ 1 + f^{(1)} + f^{(2)} + f^{(3)}(\epsilon) \right] \quad (23)$$

for a 3d chaotic grain. Substituting this expression into the gap equation (3) and comparing powers of  $1/k_F L$ , we get a simple algebraic equation for  $f^{(1)}$  with the solution

$$\lambda f^{(1)} = (1 \pm 1) \frac{\mathcal{S}\pi}{4k_F V}. \quad (24)$$

It shows that for Dirichlet (-) boundary conditions, the superconducting order parameter for a chaotic 3d cavity does not have mesoscopic deviations of order  $1/k_F L$ . This suppression is a hallmark of the chaotic case and appears due to the fluctuations of the interaction matrix elements. It can be also found by substituting  $\lambda \rightarrow \lambda(1 + \mathcal{S}\pi/4k_F V)$  into Eq. (5), which accounts only for the surface contribution to the density of states, and expanding the modified  $\Delta_0$  to first order in  $1/k_F L$  [21].

The second order correction reads

$$\lambda f^{(2)} = \frac{2\mathcal{C}}{k_F^2 V} + 2 \left( \mp 1 + \frac{1 \pm 1}{\lambda} \right) \left( \frac{\pi \mathcal{S}}{4k_F V} \right)^2 + \tilde{g}(0), \quad (25)$$

with

$$\tilde{g}(0) = \frac{2\pi}{k_F^2 V} \sum_p A_p W(L_p/\xi) \cos(k_F L_p + \beta_p), \quad (26)$$

where the contribution of periodic orbits  $L_p$  longer than the coherence length  $\xi$  is exponentially suppressed.

Equating terms of order  $(k_FL)^{-3}$ , we obtain for  $f^{(3)}(\epsilon)$  an integral equation of the form  $f^{(3)}(\epsilon) = h(\epsilon) + \int K(\epsilon')f^{(3)}(\epsilon')d\epsilon'$ , which is solved with the ansatz  $f^{(3)}(\epsilon) = h(\epsilon) + c$ , where  $c$  is a constant. We obtain

$$f^{(3)}(\epsilon) = \frac{\pi\lambda\delta}{\Delta_0} \left[ \frac{\Delta_0}{\sqrt{\epsilon^2 + \Delta_0^2}} + \frac{\pi}{4} \right]. \quad (27)$$

Note that a) since  $\delta/\Delta_0 \ll 1$  is an additional small parameter the contribution (27) can be comparable to lower orders in the expansion in  $1/k_FL$  and b) the order parameter  $\Delta(\epsilon)$  has a maximum at the Fermi energy ( $\epsilon = 0$ ) and decreases on an energy scale  $\epsilon \sim \Delta_0$  as one moves away from the Fermi level. One can also show that mesoscopic corrections given by Eqs. (24), (25) and (27) always enhance  $\Delta(0)$  as compared to the bulk value  $\Delta_0$ . A couple remarks are in order: a) the energy dependence of the gap is universal in the sense that it does not depend on specific grain details, b) the matrix elements  $I(\epsilon, \epsilon')$  play a crucial role, e.g. they are responsible for most of the deviation from the bulk limit. Finally we briefly address the interplay of mesoscopic fluctuations and parity effects (see [28] for a more detailed account). The Matveev-Larkin (ML) parity parameter  $\Delta_p$  [23], a experimentally accessible observable, accounts for even-odd asymmetries in ultrasmall superconductors. While the ML parameter coincides with the standard superconducting gap in the bulk limit, in [23] it was found that its leading finite size correction is given by

$$\Delta_p \equiv E_{2N+1} - \frac{1}{2}(E_{2N} + E_{2N+2}) = \Delta(0) - \frac{\delta}{2}, \quad (28)$$

where  $E_N$  is the ground state energy for a superconducting grain with  $N$  electrons.

We see that these corrections to the BCS mean-field approximation are comparable to mesoscopic fluctuations but have an opposite sign. For Al it seems that mesoscopic corrections are larger than those coming from (28).

### C. 2d chaotic cavities

In this section we study a 2d superconducting chaotic grain of area  $A$ , perimeter  $\mathcal{L}$ , and linear size  $L = \sqrt{A}$ . Our starting point is the gap equation (3) together with the semiclassical expressions for the spectral density, Eqs. (7) and (8), and the matrix elements,  $I(\epsilon, \epsilon')$ , Eqs. (9-13), namely

$$I(\epsilon, \epsilon') = \frac{\lambda}{A} \left[ 1 + \frac{\mathcal{L}}{k_FA} \left[ C' + \frac{\text{Si}(4k_FL)}{\pi} \right] + \frac{\mathcal{L}}{2\pi k_FA} \left[ \text{Ci} \left( \frac{4(\epsilon - \epsilon')k_FL}{\epsilon_F} \right) - \text{Ci} \left( \frac{2(\epsilon - \epsilon')}{\epsilon_F} \right) \right] + \Pi_l \left( \frac{\epsilon - \epsilon'}{\epsilon_F} \right) \right], \quad (29)$$

where  $C' \approx 0.339\dots$  and  $\text{Si}(x), \text{Ci}(x)$  are the sine and cosine integral functions, respectively. For  $l \gg L$ , the

chaotic classical dynamics leads to a universal form for the function  $\Pi_l(w)$ ,

$$\Pi_l(w) = \frac{4}{k_F^2} \frac{\sin(k_F l w)}{w}. \quad (30)$$

As in the 3d case, the energy dependence of matrix elements implies that the equations to be solved for the gap are of integral type, and that the gap itself is energy dependent. However, unlike the 3d case, we have logarithmic corrections coming from the contribution of the matrix elements. Based on the expansion in powers of  $1/k_FL$  of the spectral density and  $I(\epsilon, \epsilon')$  [see also Eqs. (A.30) and (A.36)] we propose for a 2d chaotic grain the expansion

$$\Delta(\epsilon) = \Delta_0 \left[ 1 + f^{(\log)} + f^{(1)} + \pi^{-1} f^{(2)}(\epsilon) \right]. \quad (31)$$

Following the same steps to solve the gap equation as in the 3d case, we get to leading order,

$$\lambda f^{(\log)} = \frac{\mathcal{L} \log 2k_FL}{2\pi k_FA}. \quad (32)$$

Similar logarithmic corrections to residual interactions in 2d chaotic quantum dots in the Coulomb Blockade regime were reported in Ref. 30.

The next order correction is given by

$$\lambda f^{(1)} = (C' \pm 1) \frac{\mathcal{L}}{2k_FA} + \tilde{g}(0), \quad (33)$$

with  $(-)$  for Dirichlet and  $(+)$  for Neumann boundary conditions, respectively. The truncated spectral density  $\tilde{g}(0)$  is defined as in the 3d case, with semiclassical amplitudes corresponding to 2d systems.

Finally, the energy dependent correction to the gap in 2d chaotic grains,  $f^{(2)}(\epsilon)$  is given by the same function (27) as in 3d cavities.

We note that a) in 2d the leading finite size contribution comes from the interaction matrix elements, not from the spectral density, b) finite size effects are stronger than in 3d and the leading correction does not vanish for any boundary condition, c) since effectively there are two expansion parameters  $1/k_FL \ll 1$  – assuring the validity of the semiclassical approximation – and  $\delta/\Delta_0 < 1$  – in order to apply the BCS formalism – it can happen that in a certain range of parameters the contribution  $f^{(2)}(\epsilon)$  is dominant.

In Fig. 2 we plot the gap as a function of the energy in units of the bulk gap  $\Delta_0$  for Al grains ( $k_F \approx 17.5\text{nm}^{-1}$ ,  $\lambda \approx 0.18$ , and  $\delta \approx 7279/N$  meV where  $N$  is the number of particles), of different sizes  $L$ . Note the single peak at the Fermi energy. For the smallest grains the leading contribution is  $f^{(2)}(\epsilon)$ . This is yet another indication that the matrix elements play a dominant role in the finite size effects in superconducting metallic grains.

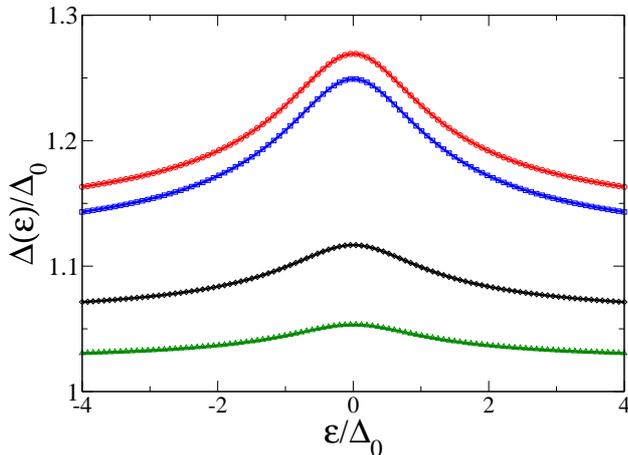


FIG. 2: Superconducting order parameter  $\Delta(\epsilon)$ , Eq. (31), in units of the bulk gap  $\Delta_0$  for 2d chaotic Al grains ( $k_F = 17.5 \text{ nm}^{-1}$ ,  $\delta = 7279/N$ ,  $\Delta_0 \approx 0.24 \text{ meV}$ ) as a function of the energy  $\epsilon$  with respect to the Fermi level  $\epsilon = 0$ . Different curves correspond to grain sizes (top to bottom) and boundary conditions:  $L = 6 \text{ nm}$ ,  $k_FL = 105$ ,  $\delta/\Delta_0 = 0.77$  (Dirichlet and Neumann boundary conditions),  $L = 8 \text{ nm}$ ,  $k_FL = 140$ ,  $\delta/\Delta_0 = 0.32$  (Dirichlet), and  $L = 10 \text{ nm}$ ,  $k_FL = 175$ ,  $\delta/\Delta_0 = 0.08$  (Dirichlet). The leading contribution comes from the energy dependent matrix elements  $I(\epsilon, \epsilon')$ .

#### IV. ENHANCEMENT OF SUPERCONDUCTIVITY IN NANOGRAINS: IDEAL VERSUS REAL GRAINS

According to the findings of previous sections the superconducting gap is an oscillating function of the system size and the number of electrons inside the grain. Even for grains with  $N \sim 10^4 - 10^5$  electrons considerable deviations from the bulk limit are observed. For a fixed grain size, the deviations from the bulk limit are the larger the more symmetric the grain is. This is a typical shell effect similar to that found in other fermionic systems, such as nuclei and atomic clusters [1]. These shell effects have their origin in the geometrical symmetries of the grain. Symmetries induce degeneracies in the spectrum and, consequently, stronger fluctuations in the spectral density. The superconducting gap is enhanced if the Fermi energy is in a region of level bunching (large spectral density). Likewise, if the Fermi energy is close to a shell closure (small spectral density) the superconducting gap will be much smaller than in the bulk limit.

Therefore, thanks to shell effects, one can adjust the gap value by adding or removing few electrons in such a way that the Fermi energy moves into a region of high or low spectral density. In fact, shell effects in metallic grains of different geometries have recently attracted considerable attention [14, 15, 17–19, 36, 37]. A superconducting spherical shell and a rectangular grain were studied numerically in Ref. 15. A similar analysis was carried out in Ref. 14 for a nanowire. A qualitative analy-

sis of a spherical superconductor was reported in Ref. 17.

Discrepancies with experiments are expected because factors such as decoherence, deformations of the shape of the grain, and surface vibrational modes are not taken into account in the theoretical analysis. In this section we discuss the impact of small deformations of the grain and of decoherence effects that shorten the coherence length. We will see that weakly deformed grains can be modeled as symmetric ones but with an effective coherence length that incorporates the details of the deformation. The semiclassical formalism utilized in this paper is especially suited to tackle this problem.

#### A. Superconductivity and shell effects

We study the dependence of the gap on the number of electrons  $N$  inside the grain and compare the gap between two grains with slightly different degree of symmetry. We focus on 3d rectangular grains where deviations from the bulk results are expected to be larger.

Since the matrix elements are constant the gap is only a function of the Fermi energy. We first compute the Fermi energy as a function of  $N$  by inverting the relation

$$\frac{1}{2}N(\epsilon_F) = \int^{\epsilon_F} g(\epsilon)d\epsilon. \quad (34)$$

Then we solve the homogeneous gap equation (3) following the steps of section III with the spectral density given by Eqs. (A.7,A.10,A.8).

The spectral density depends on the cutoff, namely, on the number of periodic orbits taken into account. This cutoff is set by the single-particle coherence length  $l$ . Here we take  $l \sim 12L$  where  $L$  is the length of the longest side of the parallelepiped and study the differences between a cubic and a rectangular grain. The cutoff is chosen to be much larger than the system size in order to observe fluctuations but considerably smaller than the superconducting coherence length  $\xi$  in order to accommodate other effects (see below) that reduce the typical single-particle coherence length in realistic nanograins.

We study a range of  $N$  such that the BCS theory is still applicable but deviations from the bulk limit are sizable. For Al grains this corresponds to  $N > 20000$ .

In Fig. 3 we plot  $\Delta$ , Eq.(14), as a function of  $N$ . The sides  $a_1 : a_2 : a_3$  were chosen to be 2.1 : 1.9 : 1 (parallelepiped) and 1.59 : 1.59 : 1.59 (cube). For both settings we observe strong fluctuations with respect to the bulk value. The fluctuations are clearly stronger in the cubic case since the grain symmetry is larger. We also observe that a slight modification of the grain size (or equivalently  $N$ ) can result in substantial changes in the gap value, see Fig. 3.

The magnitude of these fluctuations is in qualitative agreement with the theoretical prediction,  $\frac{\tilde{\Delta}}{\Delta_0} \approx \sqrt{\frac{\pi\delta}{4\Delta_0}}$  [20]. The observed differences between the cube and the parallelepiped are due to the different symmetry of these

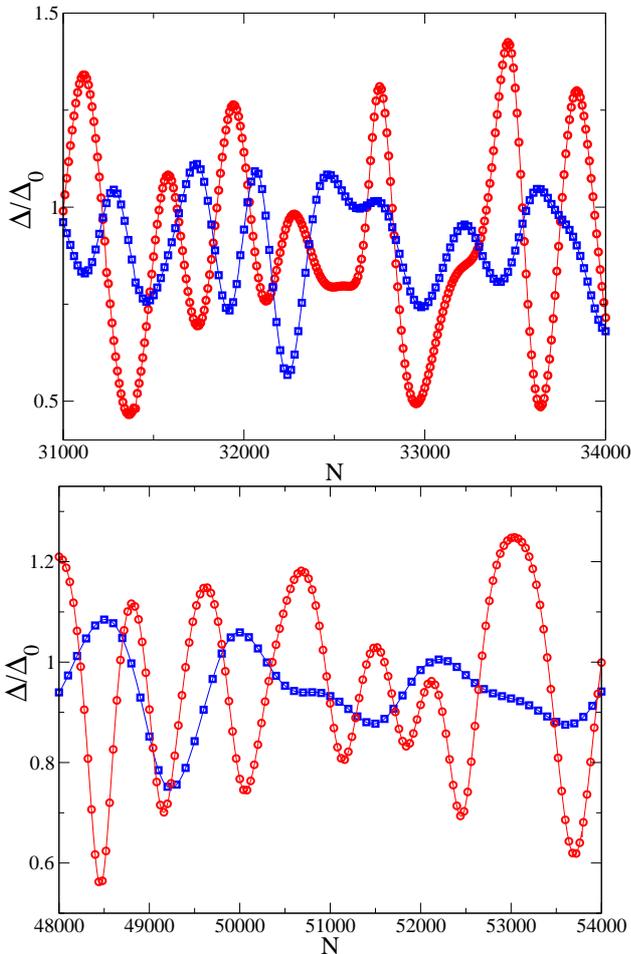


FIG. 3: The superconducting gap  $\Delta$ , Eq.(14) in units of  $\Delta_0$ , for a cubic (circles) and parallelepiped-shaped (2.1 : 1.9 : 1) (squares) superconducting grain of volume  $N/181 \text{ nm}^3$  versus the number of electrons  $N$  inside the grain. Fluctuations are on average stronger in the cubic grain due to its larger symmetry. The parameters utilized are those corresponding to superconducting Al grains.

grains. In the cube the overall symmetry factor in the spectral density is  $\propto N^{1/2}$ . The parallelepiped has only two symmetry axis and therefore the symmetry factor  $\sim N^{1/3}$ . We expect cube fluctuations to be stronger by a factor  $\sim N^{1/6}$ .

In addition to the fluctuations due to periodic orbits, we also expect smooth corrections to the bulk limit due to the surface and perimeter term of the spectral density. These corrections will be clearly observed as the coherence length is shortened and the contribution of periodic orbits is therefore suppressed.

### B. Finite size effects in real small grains

Highly symmetric shapes are hard to produce in the laboratory. It is thus natural to investigate to what

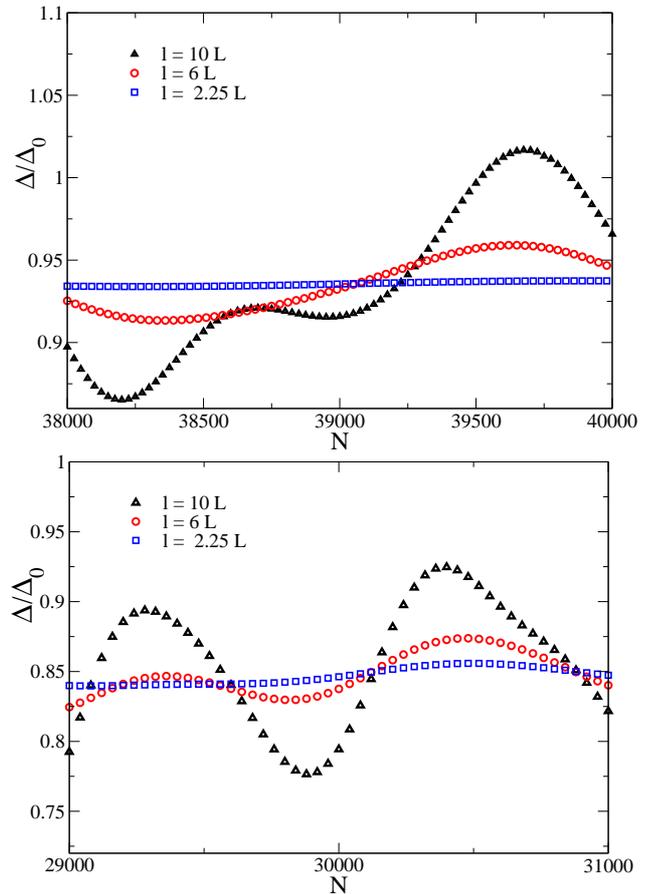


FIG. 4: Superconducting gap  $\Delta$  Eq.(14) for a cubic grain (volume  $N/181 \text{ nm}^3$ ) for different coherence lengths  $l = 2.25L$ ,  $l = 6L$ ,  $l = 10L$  in units of  $\Delta_0$  as a function of the number of particles  $N$ . As the coherence length is reduced less periodic orbits contribute to the spectral density and fluctuations are smaller. Fluctuations are strongly suppressed for coherence lengths  $l \leq 2L$ . In this limit the gap is still smaller than  $\Delta_0$  as a consequence of the surface and curvature terms in Eq.(14).

extent small deformations from a perfect cubic shape weaken the finite size effects described in previous sections. For applications it is also important to understand the dependence of the results on the single particle coherence length  $l$ . In order to study this dependence, we assume that the superconducting coherence length  $\xi$  is the largest length scale in the system. This is the most interesting region because in the opposite case  $l \gg \xi$  the results for the gap (19) are to a great extent independent of  $l$ . By contrast, in the limit  $\xi \gg l$ , the cutoff (18) induced by  $\xi$  has little effect as the contribution of periodic orbits  $L_p \geq \xi$  is already strongly suppressed by the cutoff induced by  $l$ . If  $l \sim \xi$  both cutoffs must be taken into account.

We now address these two related issues. We note that not only the effect of a finite coherence length  $l$  but also small deviations from symmetric shapes can be included in our analytical expressions for the gap by adding an

additional cutoff  $\mathcal{D}$  (besides Eq. (18)) which suppresses the contribution of periodic orbits longer than  $\mathcal{D}$ . The details of  $\mathcal{D}$  depend strongly on the source of decoherence or the type of weak deformation. Indeed, in certain cases  $\mathcal{D}$  may modify not only the amplitude but also the phase of the contribution of the periodic orbit to the trace formula used to compute the spectral density. For instance the effect of small multipolar corrections to an otherwise spherical grain [38] are included by adding an additional  $\mathcal{D}$  cutoff in term of a Fresnel integral that smoothly modulates the amplitude and phase of the periodic orbits in an ideal spherical grain.

If the deformation is given by additional small, non overlapping bumps, [39], the cutoff is exponential and only affects the amplitude. The numerical value of the cutoff depends on the original grain and is directly related to the typical size of the bump. If the source of decoherence is due to finite temperature effects, [40],  $\mathcal{D} = \frac{L_p/l}{\sinh(L_p/l)}$  with  $l$  inversely proportional to the temperature.

In Fig. 4 we show the effect of a finite coherence length  $l$  in the superconducting cubic grain investigated previously. The gap equation was solved exactly with the spectral density given by Eqs. (A.7,A.12,A.10). Instead of focusing on a specific type of deformation or decoherence mechanism we used a simple exponential cutoff for a qualitative estimate of the effect of reducing  $l$ .

The cutoff Eq. (18), related to the superconducting coherence length, does not affect the calculations as it is much longer ( $\sim 1600$  nm) than the ones employed in Fig. 4. Similar results are obtained if the analytical result (19) is utilized.

As expected, the amplitude is reduced and the fine structure of the fluctuations is washed out as the coherence length is shortened. In the range of  $N$  of Fig. 4 we did not observe any gap oscillations with  $N$  for  $l \geq 2.5L$ . This can be regarded as an effective threshold for a future experimental verification of shell effects in superconductivity. Non oscillatory corrections of the bulk result due to the surface  $\mathcal{S}$  (or perimeter  $\mathcal{L}$  in 2d) term in the spectral density are not affected by the coherence length and should be clearly observed in experiments. Note that  $\Delta$  in Fig. 4 is on average below the bulk value even for the maximum  $N$  investigated. This is a direct consequence of the negative sign of the surface term in the spectral density for Dirichlet boundary conditions used in the numerical calculations ( $f^{(1)}$  in Eq. (15)).

## V. CONCLUSIONS

We have determined the low energy excitation spectrum,  $E = \sqrt{\Delta(\epsilon)^2 + \epsilon^2}$  of small superconducting grains as a function of their size and shape by combining the BCS mean-field approach and semiclassical techniques. For chaotic grains the non-trivial mesoscopic corrections to the interaction matrix elements make them energy dependent, which, in turn, leads to a universal smooth en-

ergy dependence (27) of the order parameter  $\Delta(\epsilon)$ , see Fig 2. In the integrable (symmetric) case we found that small changes in the number of electrons can substantially modify the superconducting gap, see e.g. Fig 3. Due to its potential relevance for experiments, we have investigated how these shell effects decrease (Fig. 4) when the grain symmetry and/or the single-particle coherence length are reduced.

## Acknowledgments

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## Appendix: Semiclassical approximation for the density of states and the interaction matrix elements.

Semiclassical techniques such as periodic orbit theory [1] are not a common tool in the study of superconductivity however they are a key ingredient in our analytical treatment. In order to solve the gap equation Eq. (3) we first need a closed expression for the spectral density and the interaction matrix elements  $I(\epsilon, \epsilon')$ . In this Appendix we describe in detail how these quantities are computed using a semiclassical approximation for  $1/k_FL \ll 1$ , where  $k_F = k(\epsilon_F) = \frac{\sqrt{2m\epsilon_F}}{\hbar}$  is the momentum at the Fermi energy  $\epsilon_F$  and  $L$  is the linear system size. The resulting semiclassical expansion will be organized in powers (possibly fractional) of the small parameter  $1/k_FL$ .

In order to observe deviations from the bulk limit, the single-particle coherence length must be larger than the system size,  $l \geq L$ . The time scale,  $\tau \approx l/v_F$ , associated with  $l$  has a meaning of the lifetime of states near the Fermi energy. The condition  $l \geq L$  means that the Cooper pairs are composed of quasiparticles with a lifetime longer than the flight time through the system.

### 1. Density of states

We start with the analysis of the density of states. The semiclassical expression for  $\nu(\epsilon)$  for a given grain geometry is already known in the literature [1],

$$\nu(\epsilon') \simeq \nu_{\text{TF}}(0) [1 + \bar{g}(\epsilon) + \tilde{g}_l(\epsilon')] \quad (\text{A.1})$$

The spectral density gets both monotonous  $\bar{g}(\epsilon)$  and oscillating  $\tilde{g}(\epsilon)$  corrections. The monotonous correction at

the Fermi energy is given by the Weyl expansion.

$$\bar{g}(0) = \begin{cases} \pm \frac{S\pi}{4k_F V} + \frac{2\mathcal{C}}{k_F^2 V} & 3d \\ \pm \frac{\mathcal{L}}{2k_F A} & 2d \end{cases} \quad (\text{A.2})$$

for Dirichlet (−) or Neumann (+) boundary conditions. In Eq. (A.2),  $S$  is the surface area of the 3d cavity,  $\mathcal{C}$  is its mean curvature, while  $\mathcal{L}$  is the perimeter in the 2d case.

The oscillatory contribution to the density of states is sensitive to the nature of the classical motion. For a system whose classical counterpart is fully chaotic it is given to the leading order by the Gutzwiller trace formula [29],

$$\tilde{g}_l(\epsilon') = \Re \begin{cases} \frac{2\pi}{k_F^2 V} \sum_p^l A_p e^{i[k_F L_p + \beta_p]} e^{i\frac{\epsilon'}{2\epsilon_F} k_F L_p} & 3d \\ \frac{2}{k_F A} \sum_p^l A_p e^{i[k_F L_p + \beta_p]} e^{i\frac{\epsilon'}{2\epsilon_F} k_F L_p} & 2d, \end{cases} \quad (\text{A.3})$$

where we used  $k(\epsilon') \simeq k_F + \epsilon' k_F / 2\epsilon_F$ . The summation is over a set of classical periodic orbits ( $p$ ) of lengths  $L_p < l$ . Only orbits shorter than the quantum coherence length  $l$  of the single-particle problem are included. The amplitude  $A_p$  increases with the degree of symmetry of the cavity [1] (see below). In the chaotic case  $A_p = A_p(\epsilon_F)$  is given by

$$A_p(\epsilon_F) = \frac{L_p}{|\det(\mathbf{M}_p - \mathbf{I})|^{1/2}}, \quad (\text{A.4})$$

with the monodromy matrix  $\mathbf{M}_p$  taking into account the linearized classical dynamics around the periodic orbit. The classical flow also determines [1] the topological index  $\beta_p$  in Eq. A.3.

Note that Eqs (A.3) and (A.4) indicate that the scaling of  $\tilde{g}$  in terms of the small parameter

$$\zeta = 1/k_F L \quad (\text{A.5})$$

is

$$\tilde{g}_l(\epsilon') \propto \begin{cases} \zeta^2 & 3d, \\ \zeta & 2d. \end{cases} \quad (\text{A.6})$$

*Rectangular grain*

Consider a rectangular box of sides  $a_i$  with  $i = 1, \dots, d$  in  $d$  dimensions. For these systems the sum over periodic orbits is exact and given by [1, 3],

$$\tilde{g}(\epsilon') = \begin{cases} \tilde{g}^{(3)}(\epsilon') \\ -\frac{1}{2} \sum_i \sum_{j \neq i} \tilde{g}_{i,j}^{(2)}(\epsilon') \\ +\frac{1}{4} \sum_i g_i^{(1)}(\epsilon') & 3d, \\ \tilde{g}_{1,2}^{(2)}(\epsilon') \\ -\frac{1}{2} \sum_i g_i^{(1)}(\epsilon') & 2d. \end{cases} \quad (\text{A.7})$$

Here  $\tilde{g}^{(3)}$  is a sum over families of periodic orbits. Each family is parameterized by three (non simultaneously zero) integers  $\vec{n} = (n_1, n_2, n_3)$

$$\tilde{g}^{(3)}(\epsilon') = \sum_{L_{\vec{n}} \neq 0}^l j_0(k_F L_{\vec{n}} + \frac{\epsilon'}{2\epsilon_F} k_F L_{\vec{n}}) \quad (\text{A.8})$$

where  $L_{\vec{n}} = 2\sqrt{a_1^2 n_1^2 + a_2^2 n_2^2 + a_3^2 n_3^2}$  is the length of an orbit in the family and  $j_0(x) = \sin x/x$  is the spherical Bessel function. We see that

$$\tilde{g}^{(3)} \propto \zeta. \quad (\text{A.9})$$

In the same spirit,  $\tilde{g}_{i,j}^{(2)}$  is written as a sum over families of periodic orbits parallel to the plane defined by sides  $a_i, a_j$ . In this case the families are labeled by two integers  $\vec{n} = (n_1, n_2)$  and

$$\tilde{g}_{i,j}^{(2)}(\epsilon') = \begin{cases} \frac{a_i a_j \pi}{k_F V} \sum_{L_{\vec{n}} \neq 0}^l J_0 \left( k_F L_{\vec{n}}^{i,j} + \frac{\epsilon'}{2\epsilon_F} k_F L_{\vec{n}}^{i,j} \right) & 3d \\ \frac{a_i a_j}{A} \sum_{L_{\vec{n}} \neq 0}^l J_0 \left( k_F L_{\vec{n}}^{i,j} + \frac{\epsilon'}{2\epsilon_F} k_F L_{\vec{n}}^{i,j} \right) & 2d \end{cases} \quad (\text{A.10})$$

where  $L_{\vec{n}}^{i,j} = 2\sqrt{a_i^2 n_1^2 + a_j^2 n_2^2}$  is the length the orbit  $(n_1, n_2)$  and  $J_0$  is a Bessel function. Using the asymptotic expression for  $J_0$ , we find that this contribution scales with  $\zeta$  as,

$$\tilde{g}_{i,j}^{(2)}(\epsilon') \propto \begin{cases} \zeta^{3/2} & 3d \\ \zeta^{1/2} & 2d. \end{cases} \quad (\text{A.11})$$

Finally, for  $\tilde{g}_i^{(1)}$  we have periodic orbits labeled by a single integer  $n$

$$\tilde{g}_i^{(1)}(\epsilon') = \begin{cases} \frac{4\pi a_i}{k_F V} \sum_{L_n^i}^l \cos \left( k_F L_n^i + \frac{\epsilon'}{2\epsilon_F} k_F L_n^i \right) & 3d \\ \frac{4a_i}{k_F A} \sum_n \cos \left( k_F L_n^i + \frac{\epsilon'}{2\epsilon_F} k_F L_n^i \right) & 2d \end{cases} \quad (\text{A.12})$$

with lengths  $L_n^i = 2na_i$ . The dependence on  $\zeta$  in this case is

$$\tilde{g}_i^{(1)}(\epsilon') \propto \begin{cases} \zeta^2 & 3d \\ \zeta & 2d. \end{cases} \quad (\text{A.13})$$

It is important to note that depending on the classical dynamics and the spatial dimensionality there are different types of scaling with  $\zeta$ . The amplitude of the spectral fluctuations increases with the degree of symmetry of the cavity. It is maximal in spherical cavities and minimal in cavities with no symmetry axis [1]. The latter typically includes chaotic cavities, namely, cavities such that the motion of the classical counterpart is chaotic.

This relation between symmetry and fluctuations can be understood as follows. In grains with one or several symmetry axis there exist periodic orbits of the same length. As a result of taking all these degenerate orbits

into account, the amplitude of the spectral density is enhanced by a factor  $\zeta^{-1/2}$  for each symmetry axis [41, 42]. For instance, a spherical cavity has three symmetry axis so the symmetry factor is proportional to  $\zeta^{-3/2} \gg 1$ . Periodic orbits in chaotic cavities are not in general degenerate and the symmetry factor is therefore equal to one. For the range of sizes  $L \sim 5 - 10\text{nm}$  studied in this paper the difference between a chaotic and an integrable grain can be orders of magnitude.

## 2. Interaction matrix elements

### a. Semiclassical approximation to the average density

Unlike the case of the density of states, there is no general semiclassical theory for quantities, such as the interaction matrix element  $I(\epsilon, \epsilon')$ , involving the spatial integration of more than two eigenfunctions in clean systems. For integrable systems the ergodic condition,

$$I(\epsilon, \epsilon') = \frac{\lambda}{\Omega} \quad (\text{A.14})$$

with  $\Omega = V$  or  $A$  in 3d and 2d respectively, is typically not met due to the existence of constants of motion. The constraints imposed by conservation laws effectively localize the eigenfunctions in a smaller region of the available phase space.

On the other hand, for chaotic systems Eq. (A.14) is well justified as a result of the quantum ergodicity theorem [43]. The vast majority of the eigenfunctions spread almost uniformly over the whole volume (area) due to the lack of constants of motion besides the energy. If the position  $\vec{r}$  is far enough from the boundaries, we have

$$|\psi_n^2(\vec{r})|^2 = \frac{1}{\Omega}(1 + O(\zeta)) \quad (\text{A.15})$$

for almost all states close to the Fermi energy. In order to evaluate explicitly deviations from Eq. (A.14), we propose the replacement

$$|\psi_n^2(\vec{r})|^2 \rightarrow \langle |\psi(\vec{r})|^2 \rangle_{\epsilon_n}. \quad (\text{A.16})$$

The average is over a small window of states around  $\epsilon_n$ . The width of this window is controlled by an energy scale  $\hbar/\tau$  related to the single-particle coherence length  $l \approx v_F\tau$ . This averaging procedure is justified since eigenfunctions of classically chaotic systems have well defined statistical properties [45].

More specifically we define the above average as,

$$\begin{aligned} \langle |\psi(\vec{r})|^2 \rangle_{\epsilon} &= \frac{1}{g(\epsilon)} \sum_{\epsilon_n} w(\epsilon - \epsilon_n) |\psi_{\epsilon_n}(\vec{r})|^2 \\ &= \frac{1}{\pi g(\epsilon)} \int w(\epsilon') \Im G(\vec{r}, \vec{r}, \epsilon' - \epsilon + i0^+) d\epsilon' \end{aligned} \quad (\text{A.17})$$

where  $G(\vec{r}, \vec{r}, z)$  is the Green function of the non-interacting system at complex energy  $z$ ,  $w(x)$  is a normalized window function of width  $\hbar/\tau$  centered around  $x = 0$ , and  $g(\epsilon)$  is the density of states smoothed by  $w(x)$ .

Next, we employ the well known semiclassical approximation for the Green function as a sum over classical trajectories [29]. The effect of the zero-length paths joining  $\vec{r}$  with  $\vec{r}$  in zero time is to produce a constant background independent of the position (see, for example [34]) given by  $\bar{g}$  times the volume (area). The non-zero paths give the typical spatial oscillations of the average intensity. Further, using the expression for the density of states and factorizing the Thomas-Fermi density, we obtain,

$$\langle |\psi(r)|^2 \rangle_{\epsilon} = \frac{1}{\Omega} \frac{1 + \tilde{R}(\vec{r}, \epsilon)}{1 + \bar{g}(\epsilon_F) + \tilde{g}(\epsilon)} \quad (\text{A.18})$$

In both 3d and 2d,  $\tilde{R}(\vec{r}, \epsilon)$  is a sum over classical paths  $\gamma(\vec{r}) = \gamma$  starting and ending at point  $\vec{r}$  with finite lengths  $L_{\gamma}(\vec{r}) = L_{\gamma} < l$  and actions  $S_{\gamma}(\vec{r}) = \hbar k(\epsilon) L_{\gamma}$  [47]

$$\tilde{R}(\vec{r}, \epsilon) = \sum_{\gamma}^l D_{\gamma} \cos \left( k_F L_{\gamma} + \frac{\epsilon}{2\epsilon_F} k_F L_{\gamma} + \beta_{\gamma} \right). \quad (\text{A.19})$$

The classical properties of each trajectory are encoded in its topological phase  $\beta_{\gamma}$  (equal to  $\pi/4$  times the number of conjugate points reached by the trajectory) and the smooth function  $D_{\gamma} = D_{\gamma}(\vec{r}, \vec{r}', \epsilon_F)|_{\vec{r}=\vec{r}'}$  [1, 29],

$$D_{\gamma}(\vec{r}, \vec{r}', \epsilon_F) = \begin{cases} \frac{1}{k_F} \left| \det \frac{\partial^2 L_{\gamma}(\vec{r}, \vec{r}')}{\partial q_i \partial q'_j} \right|^{1/2} & \text{3d,} \\ \sqrt{\frac{2}{\pi k_F}} \left| \frac{\partial^2 L_{\gamma}(\vec{r}, \vec{r}')}{\partial q \partial q'} \right|^{1/2} & \text{2d.} \end{cases} \quad (\text{A.20})$$

Here  $q_i$  and  $q'_j$  are local coordinates transverse to the trajectory  $\gamma$  at points  $\vec{r}$ , respectively, and  $\vec{r}'$ , and  $L_{\gamma}(\vec{r}, \vec{r}')$  is its length. In 3d we have two perpendicular components, while in 2d there is only one.

Inspection of Eq. (A.20) shows that  $\tilde{R}$  scales as

$$D_{\gamma} \propto \begin{cases} \zeta & \text{3d} \\ \zeta^{1/2} & \text{2d} \end{cases}. \quad (\text{A.21})$$

Furthermore, the normalization condition implies

$$\frac{1}{\Omega} \int \tilde{R}(\vec{r}, \epsilon) d\vec{r} = \bar{g}(\epsilon) + \tilde{g}(\epsilon). \quad (\text{A.22})$$

Eq. (A.22) can also be used as the definition of the density of states without the Thomas-Fermi contribution in terms of the trace of the Green function.

The separation between smooth,  $\bar{g}(\epsilon) \simeq \bar{g}(0)$ , and oscillatory terms  $\tilde{g}(\epsilon)$  in Eq. (A.22) is as follows. Smooth contributions come from trajectories starting and ending at  $\vec{r}$  after hitting the boundary only once  $L_{\gamma} < L$ . On the other hand, trajectories hitting the boundary more than once will have in general  $L_p > L$ , and their contribution to the spatial integral can be evaluated using the stationary phase approximation to give  $\tilde{g}(\epsilon)$ .

Using Eqs. (A.2,A.6,A.21) and (A.22) the interaction matrix elements have the following semiclassical expan-

sion,

$$I(\epsilon, \epsilon') = \begin{cases} \frac{\lambda}{V} [1 + \bar{I}(\epsilon_F, \epsilon, \epsilon') - \frac{\mathcal{S}^2 \pi^2}{16k_F^2 V^2}] & \text{3d,} \\ \frac{\lambda}{A} [1 + \bar{I}(\epsilon_F, \epsilon, \epsilon') - \frac{\mathcal{L}}{2k_F A} + \tilde{g}l(\epsilon_F)] & \text{2d,} \end{cases} \quad (\text{A.23})$$

where

$$\bar{I}(\epsilon, \epsilon') = \frac{1}{\Omega} \int \tilde{R}(\vec{r}, \epsilon) \tilde{R}(\vec{r}, \epsilon') d\vec{r}. \quad (\text{A.24})$$

b. *Evaluation of  $\bar{I}(\epsilon, \epsilon')$*

As we will see,  $\bar{I}(\epsilon, \epsilon')$  is a smooth function of both  $\epsilon$  and  $\epsilon'$ ; it does not oscillate as rapidly as  $e^{iS/\hbar}$  where  $S$  is the classical action. The key point in carrying out the spatial integration in Eq. (A.24) is the separation of  $\tilde{R} = \tilde{R}^{\text{short}} + \tilde{R}^{\text{long}}$  into short and long classical trajectories. A similar separation leads to the smooth and oscillatory contributions to the density of states discussed in the previous section. In other words, our approach to evaluating  $\bar{I}(\epsilon, \epsilon')$  is similar to the Weyl expansion for the density of states.

To calculate  $\tilde{R}^{\text{short}}$ , we note that in the regime  $l \geq L$ , the short trajectories of length  $L_\gamma < L$  are insensitive to the smoothing, and hence their contribution to the imaginary part of the Green function in (A.17) can be pulled out from the energy integration. This means that  $\tilde{R}^{\text{short}}$  is simply proportional to the imaginary part of the Green function associated with the short paths.

Following Balian and Bloch [41] the basic idea of the subsequent calculation is that the boundary of the grain can be locally approximated as a plane in 3d and a straight line in 2d provided that the observation point is close enough to it. Within this approximation, the exact Green function representing a single reflection off an infinite wall can be calculated using the method of images for the Green function  $G^0$ , which describes a freely propagating particle (for the 3d case Eq. (A.20) is quantum mechanically exact and gives the same result).

Using the free Green function

$$G^0(\vec{r}, \vec{r}', \epsilon + i0^+) = \begin{cases} -\frac{m}{2\pi\hbar^2} \frac{e^{ik(\epsilon)|\vec{r}-\vec{r}'|}}{|\vec{r}-\vec{r}'|} & \text{3d} \\ -\frac{im}{2\hbar^2} H_0^+(k(\epsilon)|\vec{r}-\vec{r}'|) & \text{2d,} \end{cases} \quad (\text{A.25})$$

where  $H_0^+$  is the Hankel function, applying the method of images, and taking the imaginary part we get

$$\tilde{R}^{\text{short}}(\vec{r}, \epsilon) = \pm \begin{cases} \frac{\sin 2k(\epsilon)x}{2k(\epsilon)x} & \text{3d,} \\ J_0(2k(\epsilon)x) & \text{2d,} \end{cases} \quad (\text{A.26})$$

where  $x$  is the distance from the point  $\vec{r}$  to the boundary, and the minus (plus) sign corresponds to Dirichlet (Neumann) boundary conditions.

After  $\tilde{R}^{\text{short}}$  is plugged into Eq. (A.24), the integral along directions parallel to the plane simply yields a factor of  $\mathcal{S}$  in 3d and  $\mathcal{L}$  in 2d. The integration in the perpendicular direction is naturally truncated at the system

linear size  $L$ . In 3d, using  $\int_0^L = \int_0^\infty - \int_L^\infty$ , we obtain

$$\bar{I}_{3d}^{\text{short}}(\epsilon, \epsilon') = -\frac{\mathcal{S}}{8k_F^2 LV} + \frac{\pi\mathcal{S}}{4V} \frac{\text{Min}[k(\epsilon), k(\epsilon')]}{k(\epsilon)k(\epsilon')},$$

which, as expected, is a smooth function of  $\epsilon$  and  $\epsilon'$ . The second term in this expression was previously obtained in Ref. 21 via a slightly different method, which however misses the first term of order  $\zeta^2$ .

A similar analysis in 2d is more subtle due to divergence of the integration in the direction perpendicular to the boundary when it tends to infinity. However, there is a natural upper limit for this integration given by the linear system size  $L$ . Upon using  $k(\epsilon) \simeq k_F(1 + \epsilon/2\epsilon_F)$  and introducing the scaled perpendicular distance to the boundary  $y = 2k_F x$ ,

$$\bar{I}_{2d}^{\text{short}}(\epsilon, \epsilon') = \frac{\mathcal{L}}{2k_F A} \times \quad (\text{A.27})$$

$$\int_0^{2k_F L} J_0 \left[ \left(1 + \frac{\epsilon}{2\epsilon_F}\right) y \right] J_0 \left[ \left(1 + \frac{\epsilon'}{2\epsilon_F}\right) y \right] dy.$$

Employing the asymptotic expression for the Bessel functions, we find

$$\bar{I}_{2d}^{\text{short}}(\epsilon, \epsilon') = \frac{\mathcal{L}}{2k_F A} \times \quad (\text{A.28})$$

$$\left[ C + \frac{1}{\pi} \int_1^{2k_F L} \frac{\sin 2y + \cos 2(\epsilon - \epsilon')y/\epsilon_F}{y} dy \right]$$

valid for  $k_F L \gg 1$ . In Eq. (A.28) the constant  $C = \int_0^1 J_0^2(y) dy \simeq 0.850 \dots$ . We see that, contrary to the 3d case,  $\bar{I}_{2d}^{\text{short}}$  depends on energy (through the difference  $\epsilon - \epsilon'$ ). This implies that in 2d chaotic systems the superconducting gap is energy dependent even to leading order in  $\zeta$ .

The integrals in Eq. (A.28) can be expressed in terms of the sine-integral (Si) and cosine-integral (Ci) functions. Our final results is

$$\begin{cases} \bar{I}_{3d}^{\text{short}}(\epsilon_F) = \frac{\pi\mathcal{S}}{4k_F V} & \text{3d,} \\ \bar{I}_{2d}^{\text{short}}(\epsilon_F, \epsilon - \epsilon') = \frac{\mathcal{L}}{k_F A} \left[ C' + \frac{\text{Si}(4k_F L)}{\pi} \right] \\ + \frac{\mathcal{L}}{2\pi k_F A} \left[ \text{Ci} \left( \frac{4(\epsilon - \epsilon')k_F L}{\epsilon_F} \right) - \text{Ci} \left( \frac{2(\epsilon - \epsilon')}{\epsilon_F} \right) \right] & \text{2d} \end{cases} \quad (\text{A.29})$$

with  $C' = C - \text{Si}(2)/\pi = 0.339 \dots$ . Thus for fixed  $\epsilon$  and  $\epsilon'$ ,  $\bar{I}^{\text{short}}$  scales with  $\zeta \ll 1$  as follows

$$\begin{cases} \bar{I}_{3d}^{\text{short}}(\epsilon_F) \propto \zeta + b\zeta^2 & \text{3d,} \\ \bar{I}_{2d}^{\text{short}}(\epsilon_F, \epsilon - \epsilon') \propto \zeta + b'\zeta \log \zeta & \text{2d,} \end{cases} \quad (\text{A.30})$$

where  $b$  and  $b'$  are constants independent of the system size. Note the non-algebraic dependence on  $\zeta$  in the 2d case. The constant  $b$  turns out to be much smaller than all other second order contributions to the gap, and will be dropped from now on.

Now we focus on the contribution of long paths,  $\tilde{R}^{\text{long}}$ , to the spatial integral (A.24). As is standard in these cases, we evaluate the smooth functions  $D_\gamma$  in  $\tilde{R}^{\text{long}}$  at the Fermi energy and expand  $k(\epsilon) \simeq k_F + k(\epsilon)^2/2k_F$ . The phases involved in the spatial integration have the form

$$\begin{aligned} \Phi_{\gamma,\gamma'}^\pm(\epsilon, \epsilon', \vec{r}) &= k_F(L_\gamma \pm L_{\gamma'}) \\ &+ \frac{k_F}{2\epsilon_F}(L_\gamma \epsilon \pm L_{\gamma'} \epsilon'), \end{aligned} \quad (\text{A.31})$$

where  $L_\gamma(\vec{r})$  is the length of the trajectory  $\gamma$ . In chaotic systems different trajectories, in general, will have lengths differing by at least  $L$  (see, however [48]). This means that since the first term in Eq. (A.31) scales as  $1/\zeta \gg 1$ , the integral over  $r$  in  $\Phi_{\gamma,\gamma'}^\pm(\epsilon, \epsilon', \vec{r})$  and  $\Phi_{\gamma,\gamma}^\pm(\epsilon, \epsilon', \vec{r})$  can be evaluated by the stationary phase method. Each spatial integration [1] yields an extra factor  $\propto 1/\zeta^{1/2}$ . Combining this with the prefactors (A.21), we find that the so-called non-diagonal terms  $\bar{I}_{\text{ndg}}^{\text{long}}$  are of order

$$\bar{I}_{\text{ndg}}^{\text{long}}(\epsilon, \epsilon') \propto \begin{cases} \zeta^{5/2} & \text{3d,} \\ \zeta^2 & \text{2d.} \end{cases} \quad (\text{A.32})$$

On the other hand, terms that involve  $\Phi_{\gamma,\gamma}^-(\epsilon, \epsilon', \vec{r})$  do not oscillate rapidly, because in this case the phase, namely, the second term in Eq. (A.31), scales as  $\zeta$  and not as  $1/\zeta$ . This contribution involves coherent double sums over classical trajectories and is usually referred to as the diagonal contribution,  $\bar{I}_{\text{dg}}^{\text{long}}$ . It can be expressed in terms of a purely classical quantity

$$\Pi_l(w) = \int \sum_\gamma^l D_\gamma^2 \cos w k_F L_\gamma(\vec{r}) d\vec{r} \quad (\text{A.33})$$

as follows

$$\bar{I}_{\text{dg}}^{\text{long}}(\epsilon_F, \epsilon - \epsilon') = \begin{cases} \frac{1}{V} \Pi_l \left( \frac{\epsilon - \epsilon'}{\epsilon_F} \right) & \text{3d} \\ \frac{1}{A} \Pi_l \left( \frac{\epsilon - \epsilon'}{\epsilon_F} \right) & \text{2d.} \end{cases} \quad (\text{A.34})$$

Keeping also in mind the  $\zeta$ -dependence of the coefficients  $D_\gamma$ , we have

$$\bar{I}_{\text{dg}}^{\text{long}}(\epsilon_F, \epsilon - \epsilon') \propto \begin{cases} \zeta^2 & \text{3d,} \\ \zeta & \text{2d.} \end{cases} \quad (\text{A.35})$$

Equations (A.29,A.32) and (A.34) complete the evaluation of  $\bar{I}$ . Restricting ourselves to the first two orders in  $\zeta$  ( $\zeta$  and  $\zeta \log \zeta$  in the 2d case), we finally obtain

$$I(\epsilon, \epsilon') = \begin{cases} \frac{\lambda}{V} \left[ 1 + \bar{I}_{3d}^{\text{short}}(\epsilon_F) - \frac{\pi^2 S^2}{16k_F^2 V^2} + \bar{I}_{\text{dg}}^{\text{long}}(\epsilon_F, \epsilon - \epsilon') \right] & \text{3d,} \\ \frac{\lambda}{A} \left[ 1 + \bar{I}_{2d}^{\text{short}}(\epsilon_F, \epsilon - \epsilon') + \bar{I}_{\text{dg}}^{\text{long}}(\epsilon_F, \epsilon - \epsilon') \right] & \text{2d.} \end{cases} \quad (\text{A.36})$$

Equations (A.36) together with the definitions (A.29) and (A.34) allow for the calculation of interaction matrix elements in 3d and 2d chaotic grains. In general, the explicit evaluation of  $\Pi_l(w)$  requires the precise knowledge of all classical paths up to lengths of the order of the single-particle coherence length  $l$  that have a crossing at  $\vec{r}$  for every point inside the cavity. However, if  $l$  is large enough compared to  $L$  (in practice  $l \simeq 5L$  suffices) ergodic arguments can be invoked and a closed expression for the interaction matrix elements can be found. In situations when  $l \simeq L$  one must carry out the explicit system-dependent calculation.

Classical ergodicity of chaotic systems can be formulated in various ways [49], but it always implies that sums over paths with the prefactors  $D_\gamma^2$  behave as uniform distributions of the lengths of the classical paths. In the present context classical ergodicity leads to the following *sum rule* [50] for classical closed orbits:

$$\sum_{\gamma}^{l \gg L} D_\gamma^2 \delta(l - L_\gamma(\vec{r})) = \begin{cases} \frac{4\pi^2}{k_F^2 V} & \text{3d} \\ \frac{4}{k_F A} & \text{2d,} \end{cases} \quad (\text{A.37})$$

Integration over lengths up to  $l$  on both sides with a smooth weight function is also assumed. Using this result and noting that the right hand side of Eq. (A.37) is independent of the position  $\vec{r}$ , we get

$$\Pi_{l \gg L}(w) = \begin{cases} \frac{4\pi^2}{k_F^3} \frac{\sin w k_F l}{w} & \text{3d,} \\ \frac{4}{k_F^2} \frac{\sin w k_F l}{w} & \text{2d.} \end{cases} \quad (\text{A.38})$$

In the ergodic regime,  $l \gg L$ , these results enable us to evaluate explicitly the energy dependence of the interaction matrix elements in chaotic cavities.

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- [1] M. Brack and R.K. Bhaduri, *Semiclassical Physics*, (Addison-Wesley, New York 1997).  
[2] R. Balian and C. Bloch, Ann. Phys. **60**, 401 (1970).  
[3] R. Balian and C. Bloch, Ann. Phys. **64**, 271 (1971).  
[4] P. W. Anderson, J. Phys. Chem. Solids. **11**, 26 (1959).  
[5] R. Parmenter, Phys. Rev. **166**, 392 (1968).  
[6] J. M. Blatt and C. J. Thompson, Phys. Rev. Lett. **10**,

- 332 (1963); C. J. Thompson and J. M. Blatt, Phys. Lett. **5**, 6 (1963).  
[7] B. Mühlischlegel, et al, Phys. Rev. **B 6**, 1767 (1972).  
[8] J. Bardeen, L.N. Cooper, and J.R. Schrieffer, Phys. Rev. **108**, 1175 (1957).  
[9] D.C. Ralph, C.T. Black, and M. Tinkham, Phys. Rev. Lett. **74**, 3241 (1995); C.T. Black, D.C. Ralph, and M.

- Tinkham, Phys. Rev. Lett. **76**, 688 (1996).
- [10] Y. Guo, Y-F Zhang, X-Y Bao, et al., Science **306**, 1915 (2004).
- [11] W. H. Li, C. C. Yang, F. C. Tsao, and K. C. Lee, Phys. Rev. B **68**, 184507 (2003); W. H. Li, C. C. Yang, F. C. Tsao, S. Y. Wu, P. J. Huang, M. K. Chung, and Y. D. Yao, *ibid.* **72**, 214516 (2005); S. Bose, P. Raychaudhuri, R. Banerjee and P. Ayyub, Phys. Rev. B **74**, 224502 (2006); S. Reich, G. Leitus, R. Popovitz-Biro, and M. Schechter, Phys. Rev. Lett. **91**, 147001 (2003).
- [12] Mastellone A, Falci G, Fazio R, Phys. Rev. Lett. **80** 4542 (1998); Di Lorenzo A, Fazio R, Hekking FWJ, et al., Phys. Rev. Lett. **84** 550 (2000).
- [13] H. Heiselberg, Phys. Rev. A **68**, 053616 (2003).
- [14] A. A. Shanenko, et al., Phys. Rev. **B 74**, 052502 (2006); M. D. Croitoru, et. al, Phys. Rev. **B 76**, 024511 (2007).
- [15] V.N. Gladilin, et al., Solid State Comm. **121**, 519 (2002).
- [16] R.W. Richardson, Phys. Lett. **3**, 277 (1963); R.W. Richardson and N. Sherman, Nucl. Phys. **B 52**, 221 (1964).
- [17] H.Boyaci, Z.Gedik, I.Kulik, J. of Supercond. **14** ,133 (2001); V.N. Gladilin, J. Tempere, I.F. Silvera, and J.T. Devreese, Phys. Rev. **B 74**, 104512 (2006);J. Tempere, V.N. Gladilin, I.F. Silvera and J.T. Devreese, Phys. Rev. **B 72**, 094506 (2005).
- [18] V. Z. Kresin and Y. N. Ovchinnikov, Phys. Rev. **B 74**, 024514 (2006); Y.N. Ovchinnikov and V.Z.Kresin, Eur. Phys. J. **B45**, 5 (2005).
- [19] V. Kresin, Y. N. Ovchinikov, and S. A. Wolf, Phys. Rep., **431**,231 (2006); B. Cao, C. M. Neal, A. K. Starace et al., arXiv:0804.0824.
- [20] H. Olofsson, S. Aberg, and P. Leboeuf, Phys. Rev. Lett. **100**, 037005 (2008).
- [21] M. Farine, et al. , Phys. Rev. B **68**, 024507 (2003).
- [22] J. Dukelsky, et al., Phys. Rev. Lett. **83**, 172 (1999); Phys. Rev. **B 61**, 12302 (2000); J.M. Roman, et al., Nucl. Phys. **B634**, 483 [FS] (2002); R. A. Smith et al., Phys. Rev. Lett. **77**, 4962 (1996).
- [23] K. A. Matveev and A. I. Larkin, Phys. Rev. Lett. **78**, 3749 (1997).
- [24] J.M. Gordon, et al., Phys. Rev. B **35**, 4909 (1987).
- [25] M. Ma and P. A. Lee, Phys. Rev. B **32**, 5658 (1985).
- [26] M. Schechter, et al., Phys. Rev. B. **63**, 214518 (2001).
- [27] E. A. Yuzbashyan, et al., Phys. Rev. B. **71**, 094505 (2005).
- [28] A. M. Garcia-Garcia, J. D. Urbina, E. A. Yuzbashyan, K. Richter, and B. L. Altshuler, Phys. Rev. Lett. **100**, 187001 (2008).
- [29] M. Gutzwiller, *Chaos in Classical and Quantum Mechanics* (Springer, New York, 1990).
- [30] See for example S. Tomsovic, D. Ullmo, and A. Bäcker, Phys. Rev. Lett. **100**, 164101 (2008) and references therein.
- [31] A. Goshal, M. Randeria, and N. Trivelpid, Phys. Rev. B. **65**, 014501 (2001).
- [32] B. Janko, et al., Phys. Rev. **B 50**, 1152 (1994).
- [33] M.T. Tuominen, J.M. Hergenrother, T.S. Tighe, and M. Tinkham, Phys. Rev. Lett. **69**, 1997 (1992).
- [34] E. B. Bogomolny, Physica (Amsterdam) **31D**, 169 (1988).
- [35] M. Berry, J. Phys. A **10**, 2083 (1977).
- [36] R. Moro, X. Xu, S. Yin, and W. de Heer, Science **300**, 1265 (2003).
- [37] J. Friedel, J.Phys. **2**, 959 (1992).
- [38] P. Meier, M. Brack and S.C. Creagh, Z. Phys. D **41**, 281 (1997).
- [39] N. Pavloff and C. Schmit, Phys. Rev. **B 58**, 4942 (1998).
- [40] P. Leboeuf, VIII Hispalensis International Summer School, to appear in Lecture Notes in Physics, Springer-Verlag, Eds. J.M. Arias and M. Lozano; nucl-th/0406064.
- [41] R.Balian and B. Duplantier, Ann. Phys. **104**, 300 (1977).
- [42] R. Balian and C. Bloch, Ann. Phys. **69**, 76 (1971).
- [43] Roughly, the quantum ergodicity theorem [44] says that for classical systems with ergodic classical dynamics, the eigenfunctions converge to an uniformly distributed density in the semiclassical limit.
- [44] Y. Colin de Verdiere in *Les Houches School Lecture Series LII*, p. 251, eds. M.J. Giannoni, A. Voros and J. Zinn-Justin, (North-Holland, Amsterdam 1989).
- [45] The fact that quantum systems whose classical counterpart is chaotic have strong ergodic properties (in the sense that spatial averages for a single wavefunction are equivalent to energy averages over a set of eigenfunctions) was first conjectured by Berry [35]. The equivalent statement for disordered systems can be found, for example, in I. V. Gornyi and A.D Mirlin, Phys. Rev. E **65**, 025202 (2002). There is strong evidence, both numerical and experimental, for this conjectures, see e.g. [46].
- [46] H.-J. Stöckmann, *Quantum chaos: An introduction*, (Cambridge University Press, 1999).
- [47] J. D. Urbina and K. Richter, Phys. Rev. E **70**, 015201(R) (2004).
- [48] In certain situations, pairs of trajectories with arbitrarily small action differences can be constructed that describe important effects beyond the diagonal approximation (see for example M. Sieber and K. Richter, Physica Scripta **128** (2001)).
- [49] A. M. Ozorio De Almeida, *Hamiltonian Systems: Chaos and Quantization*, (Cambridge University Press 1988).
- [50] M. Sieber, J. Phys. A: Math. Gen. **32**, 7679 (1999).