Dirac Landau levels for surfaces with constant negative curvature

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Studies of the formation of Landau levels based on the Schrödinger equation for electrons constrained to curved surfaces have a long history. These include as prime examples surfaces with constant positive and negative curvature, the sphere [1] and the pseudosphere [2]. Now, topological insulators, hosting Dirac-type surface states, provide a unique platform to experimentally examine such quantum Hall physics in curved space. Hence, extending previous work we consider solutions of the Dirac equation for the pseudosphere for both, the case of an overall perpendicular magnetic field and a homogeneous coaxial, thereby locally varying, magnetic field. For both magnetic-field configurations, we provide analytical solutions for spectra and eigenstates. For the experimentally relevant case of a coaxial magnetic field we find that the Landau levels split and show a peculiar scaling $\propto B^{1/4}$, thereby characteristically differing from the usual linear B and $B^{1/2}$ dependence of the planar Schrödinger and Dirac case, respectively. We compare our analytical findings to numerical results that we also extend to the case of the Minding surface.

I. INTRODUCTION

Topological classification of different states of matter rapidly evolved from a rare curiosity [3, 4] into one of the hottest fields in physics [5–9]. Nowadays, topology serves as an essential tool [10–13] to characterize the ground state properties of interacting and non-interacting condensed matter-based systems – comprising the quantum spin Hall Effect [14–17], topological insulators [18–20], 3D semi-metals [21–23], topological superconductors [24], to name a few.

On the other hand, there is also geometry entering into a local description of the dynamics. In the presence of a non-trivial metric or an emergent gauge field either in real or momentum space – the underlying geometrical structure shows up in parallel transport. Covariant derivatives thus enter the stage and play a decisive conceptual role. In single-particle non-relativistic quantum mechanics, the non-trivial geometry emerges when considering Bloch states of lattice-periodic Hamiltonians that give rise to a geometrical object known as the "Fubini-Study-Berry connection" [25]. The latter is a silent eminence, for example, behind the popular family of various quantum Hall effects, including integer [26], spin [14–17] and anomalous [27] ones. For concreteness, semiclassical k-space transport of a wave packet probes the presence of a local geometrical quantity reflected in the anomalous velocity [28]. The latter stems from nontrivial parallel transport itself, due to the emergent gauge field triggered by a coupling of the wave-packet center of mass momentum with underlying magnetic, spin-orbit or exchange-field-generated textures. Complementary, global quantities, i.e. objects integrated over k-space, such as conductances, polarizabilities or other response functions, reflect the underlying topological features, as first demonstrated by the celebrated TKNN formula [29]. The latter introduced concepts like Chern number and topological invariants into the field of condensed matter.

Apart from non-trivial geometrical features in momentum space, there are exciting possibilities for curved geometry in real space. Positive spatial curvature is "naturally realized in the laboratory" on the surface of a sphere, for example, on a buckyball [30, 31]. Further options are offered by three-dimensional topological insulators (3DTIs) nanowires. The shape of the latter can be controlled at the nanoscale by growth [32, 33] and etching techniques [34], so that the topological Dirac-like states existing on their surfaces propagate through curved twodimensional space [35, 36]. Indeed, when immersed in a magnetic field cylindrical TI nanowires with varying radius offer the possibility of combining non-trivial real space geometry with non-trivial reciprocal space one, characterizing e.g. the (integer) quantum Hall regime. A notable consequence is that the nanowires' spectral and magnetotransport properties are governed by the subtle interplay between quantum confinement and geometry that involves Aharonov-Bohm (AB) and Berry phases [35-38].

Crucially for our purposes, the shape of such 3DTI nanowire naturally leads to a (locally) negative spatial curvature. Negative curvature is associated with interesting gravitational and cosmological phenomena, e.q. event horizon and Hawking and Unruh radiation. Its realization in condensed matter setups is challenging but possible, for example in certain meta-materials whose dielectric constant varies in a fashion effectively simulating the presence of a negatively curved metric [39–41]. It was thus possible to study analogues of the aforementioned gravitational and cosmological effects using ultra-short optical pulses [42], acoustic waves [43] and Bose–Einstein condensates [44]. Effective negative curvature was also implemented via planar networks of superconducting microwave resonators [45] which allow one to model strong interactions [46–49] on the hyperbolic background. Following up on recent works by some of us [35, 36], we instead propose shaped 3DTI nanowires as a viable experimental platform [34] to realize hyperbolic surfaces of (constant) negative curvature for Dirac-like electronic systems immersed in (strong) magnetic fields.

Generally, surfaces with a constant negative curvature in an overall perpendicular magnetic field have been studied earlier. This includes the examination of Schrödinger-type [50–56] and of Dirac-type [57, 58] Landau levels on a hyperbolic plane, Schrödinger-type Landau levels on the Poincare disc [2, 59] and Dirac-type Landau levels on a hyperboloid [60–62].

Here we study the formation of Landau levels [63] on surfaces of revolution with constant negative Gauss curvature for two prominent geometries, the pseudosphere and Minding surfaces, see Fig. 1. They both can be embedded in three-dimensional space and can even be potentially realized by means of TI nanowires with properly carved radial profiles. In view of such TI nanowires the Dirac case becomes relevant. We consider the solutions of the massless Dirac equation on the curved background and in non-trivial magnetic field configurations exploring geometrical effects stemming from the underlying spin and electromagnetic connections. In particular, we consider two configurations: (i) a magnetic field \vec{B}_{\perp} (with constant field strength B) perpendicular to the surface of revolution; (ii) a coaxial magnetic field \vec{B}_{\parallel} parallel to the rotation axis. The latter is much closer to experimental realizations, and to the best of our knowledge has not been addressed before [64]. For all geometries and field configurations considered, we calculate the corresponding eigenenergies and eigenfunctions, in most cases both analytically and numerically.

We find, among other results, that Dirac Landau levels on the pseudosphere and the Minding surface in a strong coaxial B-field show a Zeeman-type splitting. Most notably, these Landau levels exhibit a peculiar $B^{1/4}$ -scaling that parametrically differs from the common $B^{1/2}$ -dependence of the planar Dirac case, hence reflecting the effect of the Gaussian curvature in a measurable spectrum.

The paper is organized as follows: In Sec. II we review the derivation of the massless Dirac equation for surfaces of revolution in external magnetic fields—taking into account concepts of spin connection and minimal coupling. In Sec. III we apply this approach to surfaces of revolution with constant negative Gaussian curvature, namely, the pseudosphere and the Minding surface. We solve the underlying Dirac eigenvalue problems analytically in Sec. IV, and numerically in Sec. V. Apart from this we also provide a detailed implementation scheme for practical lattice discretization of the underlying Dirac problems in curved spaces. We conclude in Sec. VI, and to provide a presentation as self-contained as possible we include three technical appendices.

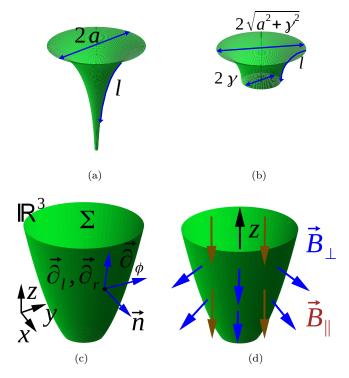


FIG. 1. Surfaces of revolution embedded in \mathbf{R}^3 : Pseudosphere (a) and Minding surface (b), including their angular and radial (arclength) parameterizations, the corresponding characteristic radii $r_{min}=\gamma$ and $r_{max}=\sqrt{a^2+\gamma^2}$ are defining outer and inner surface rims. For two magnetic configurations—the perpendicular magnetic field \vec{B}_{\perp} points along the local outer normal $\vec{n} \propto \vec{\partial}_{\phi} \times \vec{\partial}_{r}$ to the surface, while the coaxial field \vec{B}_{\parallel} points along the rotational z-axis, we set both fields as positive if $\vec{n} \cdot \vec{B}_{\perp/\parallel} > 0$, (c) Defining reference frames for a surface of revolution Σ rotated along the axial z-axis embedded in the Euclidean 3D space \mathbf{R}^3 . Vectors $\vec{\partial}_{\phi}$ and $\vec{\partial}_{r} \propto \vec{\partial}_{l}$ are tangential to Σ . The perpendicular component of the magnetic field \vec{B} points along the outer normal $\vec{n} \propto \vec{\partial}_{\phi} \times \vec{\partial}_{r}$, (d) Illustration of the two considered magnetic field configurations.

II. DIRAC EQUATION: FROM GENERAL FORMULATION TO CURVED 2D SURFACES

The formal construction of the Dirac operator for an arbitrary space-time manifold is a standard textbook material [65]. Therefore, for the sake of brevity, we just summarize the main conceptual steps and set up notation and convention. For a (D+1)-dimensional space-time manifold with a metric tensor $g_{\mu\nu}$ expressed in some local coordinates $\{x^{\mu}\}$ the corresponding Dirac equation for a massless particle reads [65, 66]

$$i\gamma^a e^\mu_a D_\mu \psi = 0 \quad , \tag{1}$$

where e_a^{μ} is the vielbein field, $g_{\mu\nu} = e_{\mu}^{a} e_{\nu}^{b} \eta_{ab}$ that brings the metric tensor $g_{\mu\nu}$ point-wisely into the canonical form $\eta_{ab} = \text{diag}(-1, +1, \dots, +1)$. Correspondingly, γ^a are the

Dirac matrices with $\{\gamma^a, \gamma^b\} = 2\eta^{ab}$. They generate the underlying Dirac (Clifford) algebra, and

$$D_{\mu} = \partial_{\mu} + \frac{1}{8} \hat{\omega}_{ab\mu} [\gamma^a, \gamma^b]$$
 (2)

stands for the covariant derivative including the spinconnection [67]

$$\hat{\omega}^{ab}_{\mu} = e^a_{\nu} \Gamma^{\nu}_{\sigma\mu} e^{\sigma b} + e^a_{\nu} \partial_{\mu} e^{\nu b}, \qquad (3)$$

see also Appendix A.

Throughout the paper we use Einstein's summation convention over repeated (one upper and one lower) Greek and Latin indices, while the metrics g and η raise and lower them, correspondingly. As usual, $\Gamma^{\nu}_{\sigma\mu}$ stand for the Christoffel symbols and $[\cdot,\cdot]$, $\{\cdot,\cdot\}$ for the associated commutators and anti-commutators, respectively. In presence of an additional electromagnetic U(1)-gauge field, described by the 1-form $A = A_{\mu}dx^{\mu}$ that couples to a charge q of the massless particle, the above covariant derivative D_{μ} changes to $D_{\mu} - i \frac{q}{\hbar} A_{\mu}$ (minimal-coupling). Later, we consider a static magnetic field \vec{B} and an electron with charge q = -e < 0.

The scalar product of two spinor fields, ψ_1 and ψ_2 , is defined as the volume integral over the D-dimensional spatial domain \mathcal{D} :

$$\langle \psi_1 | \psi_2 \rangle = \int_{\mathscr{D}} \underbrace{\sqrt{|\det g_{\text{space}}| dx^1 \cdots dx^D}}_{\omega_n} (\overline{\psi}_1)^{\mathcal{I}} (\psi_2)_{\mathcal{I}}, \quad (4)$$

where the index \mathcal{I} runs over the number of spinor-field components and the bar denotes complex conjugation. Furthermore, ω_g is an associated volume form on the spatial domain.

Particular examples of curved manifolds are 2D surfaces Σ (and wave-guides) embedded in 3D Euclidean space, $\Sigma \subset \mathbf{R}^3$. Curved space-times corresponding to them emerge by building the Cartesian product of the 2D surface with the time-coordinate line \mathbf{R} , i.e. $\mathbf{R} \times \Sigma$. In such a case, the curved space-time metric $g_{\mu\nu}$ is inherited from the flat pseudo-Euclidean metric $\langle \cdot, \cdot \rangle =$ $\operatorname{diag}(-v_F^2, +1, +1, +1)$ of the ambient space-time $\mathbf{R} \times \mathbf{R}^3$, i.e. $g_{\mu\nu} = \langle \vec{\partial}_{\mu} |_{\mathbf{R}\times\Sigma}, \vec{\partial}_{\nu} |_{\mathbf{R}\times\Sigma} \rangle$. Here the speed of light gets substituted by an effective Fermi velocity v_F as we intend to discuss fermions in metallic 2D systems. For practical computations we assume a value of $v_F =$ $5 \cdot 10^5$ m/s, which is a reasonable value considering topological insulators or Weyl-semimetals. In the 2D case $\eta_{ab} = {\rm diag}(-1,+1,+1) = e^\mu_a e^\nu_b g_{\mu\nu}$ – with the time-like component $\eta_{00} = -1$, and the space-like components $\eta_{11} = \eta_{22} = +1$, v_F becomes absorbed into the definition of the vielbein fields e_a^{μ} . Moreover, without loss of generality, we can choose the following representation of Dirac γ matrices for the given η_{ab} :

$$\gamma^0 = i\sigma_z, \, \gamma^1 = \sigma_x, \, \gamma^2 = \sigma_y \quad , \tag{5}$$

where $\sigma_{x,y,z}$ are the standard 2×2 Pauli matrices (for two spatial dimensions spinors are just two-component fields). In what follows, we keep the discussion as simple as possible and just consider 2D surfaces that emerge as surfaces of revolution, say, with the axial axis chosen along the Cartesian \hat{z} axis, see Fig. 1.

The axial symmetry makes it advantageous to choose the azimuthal angle ϕ as one coordinate of the surface. There are many convenient choices for the remaining parameter, say, radius r (if a surface is not a plain cylinder), the arclength l when walking along the surface at a fixed azimuthal angle, or the altitude z measured along the zaxis (if a surface is not flat). It turns out that the radius roffers a substantial advantage in analytical calculations, while the arclength l is better for numerical lattice simulations, as discussed in Ref. [36]. Henceforth we assume z(r) and l(r) to be well-defined functions with differentiable inverses for r ranging from r_{min} to a r_{max} – meaning we exclude cylinders, pancakes and other surfaces containing them as parts, but consider those displayed in Fig. 1. Changing parameterization from r to l is achieved by means of the relation (depending on whether $\frac{dl}{dr} \geq 0$):

$$dl = \sqrt{dx^2 + dy^2 + dz^2} = \sqrt{1 + \left(\frac{dz}{dr}\right)^2} |dr| = \pm \frac{dl}{dr} dr,$$
(6)

For a rotationally symmetric spatial domain $\Sigma \subset \mathbf{R}^3$ parameterized by (r, ϕ) via the map

$$\Sigma : [r_{min}, r_{max}] \times [0, 2\pi] \mapsto \mathbb{R}^3 (r, \phi) \to \begin{pmatrix} r \cos \phi \\ r \sin \phi \\ z(r) \end{pmatrix}, (7)$$

the spatial tangent vectors to Σ in the ambient Euclidean 3D space, see Fig. 1, are spanned by

$$\vec{\partial}_{\phi} = \frac{\partial \Sigma}{\partial \phi} = \begin{pmatrix} -r \sin(\phi) \\ r \cos(\phi) \\ 0 \end{pmatrix}, \ \vec{\partial}_{r} = \frac{\partial \Sigma}{\partial r} = \begin{pmatrix} \cos \phi \\ \sin \phi \\ \frac{dz(r)}{dr} \end{pmatrix}.$$
 (8)

Then the space-time coordinates $\{x^{\mu}\}$ on $\mathbf{R} \times \Sigma$ read (t,r,ϕ) . We define the orientation on $\mathbf{R} \times \Sigma$ by postulating $\{\vec{\partial}_{\mu}\} = \{\vec{\partial}_{t}, \vec{\partial}_{\phi}, \vec{\partial}_{r}\}$ to be the right-handed frame. Note that these vectors are orthogonal but not normalized in a sense of the ambient Pseudo-Euclidean space-time $\mathbf{R} \times \mathbf{R}^{3}$, with the flat metric $\langle \cdot, \cdot \rangle = \mathrm{diag}(-v_{F}^{2}, +1, +1, +1)$.

The induced metric tensor $g_{\mu\nu} = \langle \vec{\partial}_{\mu}, \vec{\partial}_{\nu} \rangle$ on the curved space-time $\mathbf{R} \times \Sigma$ reads

$$g_{\mu\nu} = \operatorname{diag}\left(-v_F^2, r^2, 1 + \left(\frac{dz(r)}{dr}\right)^2\right),$$
 (9)

while the vielbein field

$$e_{\mu}^{a} = \operatorname{diag}\left(v_{F}, r, \sqrt{1 + \left(\frac{dz(r)}{dr}\right)^{2}}\right)$$
 (10)

brings the metric tensor g into the corresponding Sylvester normal form $\eta=\mathrm{diag}(-1,+1,+1)$. Note that

the Greek indices (t, ϕ, r) stand for the chart-coordinates, with the corresponding coordinate vector fields $\vec{\partial}_t = \vec{\partial}_0, \vec{\partial}_\phi = \vec{\partial}_1, \vec{\partial}_r = \vec{\partial}_2$, while the Latin indices label their vielbein counterparts as captured by e^μ_a .

Computing Christoffel's symbols $\Gamma^{\sigma}_{\mu\nu}$, see Appendix B, the spin connection $\hat{\omega}^{ab}_{\mu}$, Eq. (3), yields only the non-zero entries

$$\hat{\omega}_{\phi}^{12} = -\hat{\omega}_{\phi}^{21} = \frac{1}{\sqrt{1 + \left(\frac{dz(r)}{dr}\right)^2}} = \frac{dr}{dl}.$$
 (11)

Taking all together and assuming a gauge field 1-form $A = A_{\phi}d\phi$, we obtain the following expressions for the covariant derivatives of spinors on the curved space-time $\mathbf{R} \times \Sigma$:

$$D_0 = D_t = \partial_t \,, \tag{12}$$

$$D_1 = D_{\phi} = \partial_{\phi} + \frac{i}{2}\hat{\omega}_{\phi}^{12}\sigma_z + i\frac{e}{\hbar}A_{\phi}, \qquad (13)$$

$$D_2 = D_r = \partial_r \,. \tag{14}$$

Now we can write down the final Dirac equation

$$i\gamma^a e_a^{\mu} D_{\mu} \psi = \left(-\sigma_z \frac{1}{v_F} \partial_t + i\sigma_y \, \hat{\omega}_{\phi}^{12} \, \partial_r + i\sigma_x \frac{1}{r} D_{\phi} \right) \psi = 0.$$
(15)

for massless charged fermions on $\mathbf{R} \times \Sigma$. Note that by coincidence $\hat{\omega}_{\phi}^{12} = e_2^r$.

Making use of Dirac algebra and assuming stationary states, $\psi(t,r,\phi)=e^{-iEt/\hbar}\Psi(r,\phi)$, we obtain an eigenvalue problem for the spatial eigenspinors $\Psi(r,\phi)$ in terms of the equation

$$\frac{\hat{H}}{\hbar v_F} \Psi \equiv \left(i \sigma_x \, \hat{\omega}_{\phi}^{12} \, \partial_r - i \sigma_y \frac{1}{r} D_{\phi} \right) \Psi = \frac{E}{\hbar v_F} \Psi \,. \tag{16}$$

It remains to specify A_{ϕ} for the respective configuration of the magnetic field \vec{B} . We consider two cases:

- (i) a magnetic field \vec{B}_{\perp} with a constant magnitude perpendicular to the surface of revolution Σ (in the sense of the ambient Euclidean space), and
- (ii) a homogeneous coaxial magnetic field \vec{B}_{\parallel} being aligned in parallel to the axial z-axis of the 3D ambient space, i.e. $\vec{B}_{\parallel} = (0, 0, B)$.

The Faraday tensor F at surface $\Sigma\subset {\bf R}^3$ [65] for a generic magnetic field configuration \vec{B} in 3D space takes the form

$$F = -(\vec{B} \cdot \vec{n}) \,\omega_g \tag{17}$$

where ω_g is the "volume-form" defined on Σ by the spatial part of the metric $g_{\mu\nu}$, namely

$$\omega_g = \sqrt{\det(g_{\Sigma})} \, d\phi \wedge dr = r \sqrt{1 + \left(\frac{dz(r)}{dr}\right)^2} \, d\phi \wedge dr \,.$$
 (18)

Furthermore, $(\vec{B} \cdot \vec{n})$ stands for the Euclidean scalar product of the 3D vector \vec{B} with the outer normal \vec{n} to the

surface Σ ,

$$\vec{n} = \frac{\vec{\partial}_{\phi} \times \vec{\partial}_{r}}{||\vec{\partial}_{\phi} \times \vec{\partial}_{r}||} = \frac{1}{\sqrt{1 + \left(\frac{dz}{dr}\right)^{2}}} \left(\frac{dz}{dr}\cos\phi, \frac{dz}{dr}\sin\phi, 1\right),$$
(19)

taken in the ambient space \mathbf{R}^3 . In Eq. (19) $\vec{\partial}_{\phi}$ and $\vec{\partial}_r$ are vectors in \mathbf{R}^3 , see Eq. (8), \times is the usual vector product, and $||\cdot||$ denotes the Euclidean length of a 3D vector.

Throughout the text $d\phi$ and dr are understood as differential 1-forms, and correspondingly ω_g and F, proportional to their wedge product $d\phi \wedge dr$, are 2-forms on Σ . Using the exterior derivative d, one can relate the 2-form F to the electromagnetic potential 1-form A [65]:

$$dA = F = -(\vec{B} \cdot \vec{n}) \,\omega_g \,. \tag{20}$$

As both magnetic field configurations, \vec{B}_{\perp} and \vec{B}_{\parallel} , are axially symmetric, and Σ is a surface of revolution along the axial z-axis, $(\vec{B} \cdot \vec{n})$ only depends on the radial coordinate, i.e. $(\vec{B} \cdot \vec{n})[r]$. Hence we use the symmetric electromagnetic gauge (what happens to be a Coulomb gauge) $A = A_{\phi}(r)d\phi$, where the function $A_{\phi}(r)$ satisfies

$$dA = (\partial_r A_\phi) \, dr \wedge d\phi \stackrel{!}{=} -(\vec{B} \cdot \vec{n}) \, \omega_q \,. \tag{21}$$

Its solution can be expressed as an integral

$$A_{\phi}(r) = \int_{l}^{r} (\vec{B} \cdot \vec{n})[r'] \cdot r' \cdot \sqrt{1 + \left(\frac{dz(r')}{dr'}\right)^{2}} dr' \quad (22)$$
$$= \int_{l}^{l} (\vec{B} \cdot \vec{n})[l'] \cdot r(l') dl' .$$

Explicit expressions for $A_{\phi,\perp}$ and $A_{\phi,\parallel}$ corresponding to \vec{B}_{\perp} and \vec{B}_{\parallel} are provided below.

III. THE DIRAC EQUATION ON THE PSEUDOSPHERE AND MINDING SURFACE

We apply the geometrical machinery reviewed in the previous section to surfaces of revolution that possess a constant negative Gaussian curvature

$$\kappa = -\frac{1}{a^2} \,. \tag{23}$$

Here a>0 – the radius of curvature – sets a natural spatial length scale of the problem, besides the (Fermi) wavelength. A family of surfaces of revolution Σ with a constant negative curvature κ , that are embeddable in \mathbf{R}^3 , $\Sigma:(r,\phi)\mapsto(r\cos\phi,r\sin\phi,z(r))$, are defined through the so-called *Minding equation* [68],

$$\left(\frac{dl}{dr}\right)^2 \equiv 1 + \left(\frac{dz(r)}{dr}\right)^2 = \frac{a^2}{r^2 - \gamma^2}, \qquad (24)$$

which needs to be solved for z(r) (or l(r)) that enter the above parameterization of Σ . The Minding equation has two parameters controlling the "shape" of Σ : the radius of curvature a>0 (or curvature $\kappa=-1/a^2$ itself) and the parameter $0\leq \gamma < a$ that restrains "an amount of the embeddable part" with the ambient space ${\bf R}^3$. The case $\gamma=0$ defines a pseudosphere or Beltrami surface—a surface of revolution of the tractrix—whereas the surfaces with $\gamma\neq 0$ are called Minding surfaces [68], see Fig. 1.

While the Schrödinger-type Landau levels for the pseudosphere have been considered before [69, 70], we will address in Secs. IV (analytics) and V (numerics) the case of the Dirac equation for the pseudosphere and Minding surfaces. The Dirac scenario acquires importance in view of possible experimental realizations in shaped 3DTI nanowires. Moreover, we will consider the experimentally relevant spatially homogeneous axial magnetic field, \vec{B}_{\parallel} , besides the conceptually simpler case of a constant perpendicular magnetic field, \vec{B}_{\perp} .

The embeddable part of a surface of revolution with a constant negative curvature within 3D Cartesian space only exists for a certain range of radii. Since $(dz(r)/dr)^2$ varies from zero to infinity, the middle-part of Eq. (24) becomes equal or greater than 1. Consequently, the radii r entering the right-hand side of Eq. (24) should vary only within the interval $(r_{min} = \gamma, r_{max} = \sqrt{a^2 + \gamma^2})$. [71] As mentioned above, apart from the r coordinate, also the arclength coordinate l serves convenient for the parameterization of Σ . Integrating the defining differential equation (24), one gets

$$l(r) = \pm \int_{r_0}^{r} \frac{a}{\sqrt{r'^2 - \gamma^2}} dr', \qquad (25)$$

where the \pm sign and the reference radius r_0 are free to choose. For the pseudosphere and Minding surface we set l to zero at the rim with the largest radius, $r_{max} = \sqrt{a^2 + \gamma^2}$, see Fig. 1. Then the relation

$$l_{\gamma}(r) = a \ln \left(\frac{a + \sqrt{a^2 + \gamma^2}}{r + \sqrt{r^2 - \gamma^2}} \right) \tag{26}$$

holds equally for both cases $\gamma = 0$ and $\gamma \neq 0$.

In what follows, we derive the eigenstates and eigenvalues of the Dirac Hamiltonians for massless particles moving on the pseudosphere and Minding surfaces subject to the magnetic field configurations, \vec{B}_{\perp} and \vec{B}_{\parallel} , as specified above. Analytically, we mainly focus on eigenstates that are localized well inside the surface of revolution, i.e. their probability amplitudes are falling off sufficiently fast when approaching the edges. This sets the boundary conditions of our analytical spectral problem. For the corresponding numerics, we represent a part of the surface by a discretized grid and diagonalize the Dirac Hamiltonians (with Wilson mass term [72]) using proper boundary conditions, see Sec. V.

As a>0 sets the natural length scale of the problem, we rescale all quantities with a spatial dimension with a, using the tilde symbol, thus

$$\tilde{r} = r/a$$
 , $\tilde{\gamma} = \gamma/a$ etc. (27)

Using the general results from the previous section, and Eq. (24) for $\sqrt{1 + (dz(r)/dr)^2}$, we obtain the following expressions for the spin-connection, Eq. (11), and for the corresponding electromagnetic gauge fields, Eq. (22):

$$\hat{\omega}_{\phi}^{12} = \sqrt{\tilde{r}^2 - \tilde{\gamma}^2} \,, \tag{28}$$

$$A_{\phi,\perp} = Ba^2 \sqrt{\tilde{r}^2 - \tilde{\gamma}^2} \quad , \quad A_{\phi,\parallel} = Ba^2 \frac{\tilde{r}^2}{2} \, .$$
 (29)

Furthermore, $\mathcal{E} = \hbar v_F/a$ sets the natural energy scale. Hence we normalize all quantities of dimension of energy by \mathcal{E} , but to avoid a proliferation of new symbols we do not introduce a new label for that, i.e. $H/\mathcal{E} \to H$. While for formulas and analytics we use the rescaled quantities, when plotting the spectra and effective potentials, and also for numerics we rather employ the dimensionful quantities and variables to grasp the magnitudes and units. In order to compare analytical and numerical results we use $v_F = 5 \cdot 10^5$ m/s and consider one representative pseudosphere ($\gamma = 0$) and Minding surface ($\gamma = 1 \, \text{nm}$) with the same radius of curvature $a = 60 \,\mathrm{nm}$. This choice implies that, for magnetic fields with magnitude larger than 0.15 T, the magnetic length $l_B = \sqrt{\hbar/e|B|}$ becomes shorter than the considered curvature radius.

With these rescaling conventions the Dirac Hamiltonian, Eq. (16), as a function of γ ($\gamma = 0 \rightarrow$ pseudosphere, $\gamma \neq 0 \rightarrow$ Minding surface), reads

$$\hat{H}_{\perp} = i\sigma_x \sqrt{\tilde{r}^2 - \tilde{\gamma}^2} \partial_{\tilde{r}} - i\sigma_y \frac{1}{\tilde{r}} \partial_{\phi} + \sigma_y \frac{\sqrt{\tilde{r}^2 - \tilde{\gamma}^2}}{\tilde{r}} \left(\frac{1}{\tilde{l}_B^2} + \frac{\sigma_z}{2} \right) , \qquad (30)$$

$$\hat{H}_{\parallel} = i\sigma_x \sqrt{\tilde{r}^2 - \tilde{\gamma}^2} \partial_{\tilde{r}} - i\sigma_y \frac{1}{\tilde{r}} \partial_{\phi} + \sigma_y \left(\frac{\tilde{r}}{2\tilde{l}_B^2} + \frac{\sqrt{\tilde{r}^2 - \tilde{\gamma}^2}}{\tilde{r}} \frac{\sigma_z}{2} \right).$$
(31)

Here H_{\perp}, H_{\parallel} stand for the Hamiltonian (16) respectively in the perpendicular and coaxial magnetic field configuration, while \tilde{l}_B^2 is the dimensionless magnetic area multiplied by sgn(B). Specifically

$$\frac{1}{\tilde{l}_B^2} = \operatorname{sgn}(B) \, \frac{a^2}{l_B^2} \,, \tag{32}$$

where the magnetic length

$$l_B = \sqrt{\hbar/e|B|} > 0. (33)$$

These expressions hold for both magnetic field configurations, where |B| stands for the magnitude of the $\vec{B}_{\perp/\parallel}$ field. We introduce the flux quantum $\Phi_0=h/2e$, the total area of the embeddable part of Σ , ${\rm Area}_{\Sigma}=2\pi a^2$, as well as its projection onto the plane orthogonal to its symmetry axis, ${\rm Area}_{\Sigma}'=\pi a^2=\pi (r_{max}^2-r_{min}^2).$ One obtains for \vec{B}_{\perp} and \vec{B}_{\parallel} the perpendicular and parallel fluxes $|\Phi_{\perp}|=2\pi a^2|B|$ and $|\Phi_{\parallel}|=\pi a^2|B|$, correspondingly. Thus, all together

$$\frac{1}{\tilde{l}_B^2} = \operatorname{sgn}(B) \frac{a^2}{l_B^2} = \operatorname{sgn}(B) \frac{1}{\Phi_0} \times \begin{cases} \frac{1}{2} |\Phi_{\perp}| & \text{for } \vec{B}_{\perp}, \\ |\Phi_{\parallel}| & \text{for } \vec{B}_{\parallel}. \end{cases}$$
(34)

In what follows we keep the subscripts \bot and $\|$ also for other quantities to distinguish and to trace their connections with the \vec{B}_{\bot} and $\vec{B}_{\|}$ magnetic field configurations, respectively.

It is instructive to find discrete symmetries of $\hat{H}_{\perp/\parallel}$ to simplify the spectral analysis. The fact that σ_z anticommutes with $\hat{H}_{\perp/\parallel}$ implies that the Hamiltonians have chiral symmetry:

$$\sigma_z \, \hat{H}_{\perp/\parallel} = -\hat{H}_{\perp/\parallel} \, \sigma_z \,. \tag{35}$$

As a consequence, the spectrum is symmetric with respect to E=0, and if $|\psi\rangle$ is an eigenstate of $\hat{H}_{\perp/\parallel}$ with an eigenenergy E, then $|\Xi\rangle=\sigma_z|\psi\rangle$ is also an eigenstate with eigenvalue -E. Positive and negative energy solutions are therefore directly linked by the chiral symmetry operator σ_z . Furthermore, having an eigenstate $|\psi\rangle$ of $\hat{H}_{\perp/\parallel}(\vec{B})$ with eigenenergy E for a magnetic field \vec{B} , the state $|\chi\rangle=\sigma_y\mathscr{C}|\psi\rangle$ (\mathscr{C} stands for the complex-conjugation operator) is an eigenstate of $\hat{H}_{\perp/\parallel}(-\vec{B})$ with the same energy, but for the opposite field configuration, i.e.

$$[\sigma_y \,\mathscr{C}] \, \hat{H}_{\perp/\parallel}(\vec{B}) = \hat{H}_{\perp/\parallel}(-\vec{B}) \, [\sigma_y \,\mathscr{C}] \,. \tag{36}$$

Thus, it is sufficient to examine only positively oriented magnetic fields, and for them to consider only eigenstates with non-negative energies.

IV. ANALYTICAL SOLUTIONS

A. General considerations

In the following we present analytical eigensolutions of the Hamiltonians given by Eqs. (30) and (31). For the pseudosphere we get full solutions for the perpendicular field, \vec{B}_{\perp} , and approximate solutions for the coaxial field, \vec{B}_{\parallel} , while for the Minding surfaces we only find only the zero-energy eigenstates analytically in both magnetic configurations. Numerical solutions and a cross-check of analytic results with numerics will be presented in Sec. V.

Since $\hat{H}_{\perp/\parallel}$ in Eqs. (30) and (31) do not explicitly depend on ϕ , we use the separation ansatz

$$\Psi(\phi, \tilde{r}) = \begin{pmatrix} \Psi^{\uparrow}(\tilde{r}) \\ \Psi^{\downarrow}(\tilde{r}) \end{pmatrix} e^{i(m + \frac{1}{2})\phi}, m \in \mathbb{Z}.$$
 (37)

Note that the boundary conditions of the angular part of the wavefunction are antiperiodic (see e.g. discussion in Appendix A, or Ref. [73]). This is in accordance with a local trivialization of the associated spin bundle that gives rise to the spin-connection in the form as presented by Eq. (3). For convenience we introduce effective potentials (in units of $\mathcal{E} = \hbar v_F/a$),

$$V_{\perp} = \frac{1}{\tilde{r}} \left[\left(m + \frac{1}{2} \right) + \frac{\sqrt{\tilde{r}^2 - \tilde{\gamma}^2}}{\tilde{l}_B^2} \right] , \qquad (38)$$

$$V_{\parallel} = \frac{1}{\tilde{r}} \left[\left(m + \frac{1}{2} \right) + \frac{\tilde{r}^2}{2\tilde{l}_B^2} \right], \tag{39}$$

and the auxiliary radial differential operators [74]

$$\hat{L}_{\perp/\parallel}^{\pm} = i \left[\sqrt{\tilde{r}^2 - \tilde{\gamma}^2} \left(\partial_{\tilde{r}} + \frac{1}{2\tilde{r}} \right) \pm V_{\perp/\parallel} \right]. \tag{40}$$

Using the separation ansatz, Eq. (37), and $\mathcal{E} = \hbar v_F/a$ units the eigenproblem for the Dirac Hamiltonians, Eqs. (30) and (31), takes the form

$$E\begin{pmatrix} \Psi^{\uparrow}(\tilde{r}) \\ \Psi^{\downarrow}(\tilde{r}) \end{pmatrix} = \begin{pmatrix} 0 & \hat{L}_{\perp/\parallel}^{-} \\ \hat{L}_{\perp/\parallel}^{+} & 0 \end{pmatrix} \begin{pmatrix} \Psi^{\uparrow}(\tilde{r}) \\ \Psi^{\downarrow}(\tilde{r}) \end{pmatrix} . \tag{41}$$

The effective potentials $V_{\perp/\parallel}$ that depend on the angular quantum number m act as variable mass term within the corresponding radial equations [35, 36, 72]. The radii \tilde{r} where |V| becomes minimal indicate positions where the wavefunction amplitudes, and hence the corresponding probability densities are maximal. Minima of |V| for positive (negative) magnetic fields, $\tilde{l}_B^2 > 0$ ($\tilde{l}_B^2 < 0$), require negative (non-negative) angular quantum numbers m. Physically, only minima which develop within the embeddable part of Σ can accept electrons and thus contribute to the filling factor of the Landau problem – this is in full analogy with quantum Hall states in conventional flat samples. Hence by requiring \tilde{r} to be within $\tilde{r}_{min} = \tilde{\gamma}$ and $\tilde{r}_{max} = \sqrt{1 + \tilde{\gamma}^2}$ one sets a natural cutoff for the critical value (filling factor) of m (negative for B>0 and positive for B<0):

$$m_{crit,\perp}(B) = \begin{cases} \left[-\frac{\sqrt{1-\tilde{\gamma}^2}}{\tilde{l}_B^2} - \frac{1}{2} \right], & \text{for } B > 0 \\ \left[-\frac{\sqrt{1-\tilde{\gamma}^2}}{\tilde{l}_B^2} - \frac{1}{2} \right], & \text{for } B < 0 \end{cases}$$

$$m_{crit,\parallel}(B) = \begin{cases} \left[-\frac{1}{2\tilde{l}_B^2} - \frac{1}{2} \right], & \text{for } B > 0 \\ \left[-\frac{1}{2\tilde{l}_B^2} - \frac{1}{2} \right], & \text{for } B > 0 \end{cases}$$

$$\left[-\frac{1}{2\tilde{l}_B^2} - \frac{1}{2} \right], & \text{for } B < 0 \end{cases}$$

$$(43)$$

where $\lceil \cdot \rceil$ and $\lfloor \cdot \rfloor$ are ceiling and floor functions, respectively.

We illustrate profiles of $|V_{\perp/\parallel}(\ell)|$ in the arclength coordinate, Eq. (25), for the pseudosphere in perpendicular,

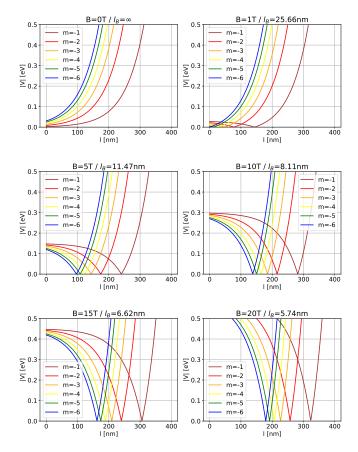


FIG. 2. Wedge-shaped effective potentials $|V_{\perp}(l)|$, Eq. (38), displayed as functions of the arclength coordinate l for a pseudosphere with $a=60\,\mathrm{nm}$ in a perpendicular magnetic field \vec{B}_{\perp} pointing along the outer normal \vec{n} , see Fig. 1(a). Different panels correspond to different strengths B, equivalently different magnetic lengths l_B . The colors code effective potentials for different angular quantum numbers m. Minima of $|V_{\perp}(l)|$ move away from the outer rim of the pseudosphere (l=0) when raising B and lowering |m|.

Fig. 2, and coaxial, Fig. 3, positive magnetic fields and different angular quantum numbers m < 0. Obviously, minima of $|V_{\perp}(l)|$ move from the outer rim (l=0) towards the cusp of the pseudosphere when raising B>0 and raising m<0. Moreover, the wedge-shaped potential wells become more pronounced upon increasing the magnetic field strengths. Contrary to this, in the absence of the field, the curves grow monotonously with l and eigenstates only exist close to the outer rim.

To solve the radial equations, Eq. (41), and compute $(\Psi^{\uparrow}(\tilde{r}), \Psi^{\downarrow}(\tilde{r}))^{\top}$ for a general E, it is convenient to decouple the underlying system of first order differential equations,

$$E \,\Psi^{\uparrow}(\tilde{r}) = \hat{L}_{\perp/\parallel}^{-} \,\Psi^{\downarrow}(\tilde{r}) \,, \tag{44}$$

$$E \Psi^{\downarrow}(\tilde{r}) = \hat{L}_{\perp/\parallel}^{+} \Psi^{\uparrow}(\tilde{r}), \qquad (45)$$

by plugging one into the other. This leads to two sepa-

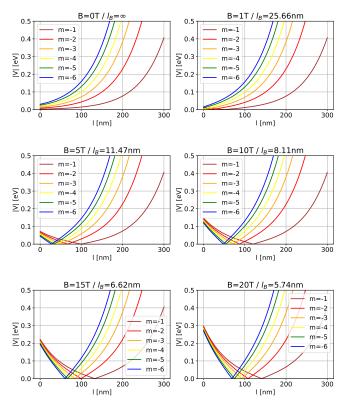


FIG. 3. Wedge-shaped effective potentials $|V_{\parallel}(l)|$, Eq. (39), displayed as functions of the arclength coordinate l, for a pseudosphere with $a=60\,\mathrm{nm}$ in a coaxial magnetic field $\vec{B}_{\parallel}=(0,0,B)$, see Fig. 1(a). Different panels correspond to different strengths B. The colors denote potentials for different angular quantum numbers m. Minima of $|V_{\parallel}(l)|$ move away from the outer rim of the pseudosphere (l=0) when raising B and lowering |m|.

rate second order differential equations for the two spinor components at eigenenergy E (in units of $\mathcal{E} = \hbar v_F/a$):

$$E^2 \Psi^{\uparrow}(\tilde{r}) = \hat{L}_{\perp/\parallel}^{-} \hat{L}_{\perp/\parallel}^{+} \Psi^{\uparrow}(\tilde{r}), \qquad (46)$$

$$E^2 \Psi^{\downarrow}(\tilde{r}) = \hat{L}_{\perp/\parallel}^+ \hat{L}_{\perp/\parallel}^- \Psi^{\downarrow}(\tilde{r}), \qquad (47)$$

that we consider below.

B. Pseudosphere

1. Zeroth Landau level

For E=0, the system of first order differential equations, Eqs. (44) and (45), decouples and one needs to solve

$$\hat{L}_{\perp/\parallel}^{+/-} \Psi^{\uparrow/\downarrow}(\tilde{r}) = 0. \tag{48}$$

In the following, the solutions of the equations containing the superscript \pm are marked with the same

symbol. Substituting $\tilde{\gamma}=0$ into the expressions for $V_{\perp/\parallel}$ and $\hat{L}_{\perp/\parallel}^{\pm}$ leads to simple radial equations with straightforward solutions. The full spinorial solutions $\left(\Psi^{\uparrow}(\tilde{r},\phi),\Psi^{\downarrow}(\tilde{r},\phi)\right)^{\top}$ for the zeroth Landau level (n=0) read:

$$\Psi_{0,\perp}^{+}(\phi,\tilde{r}) = \tilde{r}^{-\frac{1}{2} - \frac{1}{\tilde{l}_{B}^{2}}} e^{+\frac{1}{\tilde{r}}(m + \frac{1}{2}) + i\left(m + \frac{1}{2}\right)\phi} \begin{pmatrix} 1\\0 \end{pmatrix}, \quad (49)$$

$$\Psi_{0,\perp}^{-}(\phi,\tilde{r}) = \tilde{r}^{-\frac{1}{2} + \frac{1}{\tilde{l}_{B}^{2}}} e^{-\frac{1}{\tilde{r}}\left(m + \frac{1}{2}\right) + i\left(m + \frac{1}{2}\right)\phi} \begin{pmatrix} 0\\1 \end{pmatrix}. \quad (50)$$

$$\Psi_{0,\parallel}^{+}(\phi,\tilde{r}) = \tilde{r}^{-\frac{1}{2}} e^{-\frac{\tilde{r}}{2l_{B}^{2}}} e^{+\frac{1}{\tilde{r}}\left(m+\frac{1}{2}\right) + i\left(m+\frac{1}{2}\right)\phi} \begin{pmatrix} 1\\ 0 \end{pmatrix}, (51)$$

$$\Psi_{0,\parallel}^{-}(\phi,\tilde{r}) = \tilde{r}^{-\frac{1}{2}} e^{\frac{\tilde{r}}{2l_B^2}} e^{-\frac{1}{\tilde{r}}\left(m + \frac{1}{2}\right) + i\left(m + \frac{1}{2}\right)\phi} \begin{pmatrix} 0\\1 \end{pmatrix}. \tag{52}$$

To check which of them are physical we demand the amplitude of the wavefunction to vanish when $\tilde{r} \to 1$ and $\tilde{r} \to 0$.

For positive fields, solutions $\Psi^+_{0,\perp/\parallel}$ with $-1 \geq m \geq m_{crit,\perp/\parallel}$ are physically relevant, while for the negative ones, their counterparts $\Psi^-_{0,\perp/\parallel}$ with $0 \leq m \leq m_{crit,\perp/\parallel}$ span the zeroth Landau level. In any case, for a nonzero magnetic field $\vec{B}_{\perp/\parallel}$ the zeroth Dirac Landau level becomes fully spin polarized with spins pointing along ${\rm sgn}(B)\,\vec{n}$. A similar analysis was done for a 2D sphere in a homogeneous perpendicular field, see e.g. [75–77], where the zeroth Dirac Landau level is also spin polarized but electron spins are aligned along $-{\rm sgn}(B)\,\vec{n}$.

Fig. 4 shows the probability densities of representative zero-energy eigenstates in the respective effective potentials for B_{\perp} ; results for B_{\parallel} look very similar.

2. Higher Landau levels: general

To simplify the analysis, for higher Landau levels we consider just positive magnetic fields, $\tilde{l}_B^2 > 0$, as the negative ones can be obtained by means of the $\sigma_y \mathscr{C}$ symmetry stipulated by Eq. (36). We start from the second order differential equation, Eq. (46), for the up component,

$$0 = \left[\tilde{r}^2 \partial_{\tilde{r}}^2 + 2\tilde{r} \partial_{\tilde{r}} + \tilde{r} \frac{\partial V_{\perp/\parallel}}{\partial \tilde{r}} - V_{\perp/\parallel}^2 + \frac{1}{4} + E^2 \right] \Psi^{\uparrow}(\tilde{r}) \,.$$

We assume that the excited states show the same asymptotic behaviour of the radial part as the zero energy solution $\Psi^+_{0,\perp/\parallel}(\phi,\tilde{r})$, Eq. (49). This implies the following ansatz for the perpendicular and coaxial field, respectively:

$$\Psi_{\perp}^{\uparrow}(\tilde{r}) = p_{\perp}^{\uparrow}(\tilde{r})\,\tilde{r}^{-\frac{1}{2} - \frac{1}{l_B^2}}\,e^{\frac{1}{\tilde{r}}(m + \frac{1}{2})}\,,\tag{53}$$

$$\Psi_{\parallel}^{\uparrow}(\tilde{r}) = p_{\parallel}^{\uparrow}(\tilde{r}) \, \tilde{r}^{-\frac{1}{2}} \, e^{-\frac{\tilde{r}}{2l_B^2}} \, e^{\frac{1}{\tilde{r}}(m+\frac{1}{2})} \,, \tag{54}$$

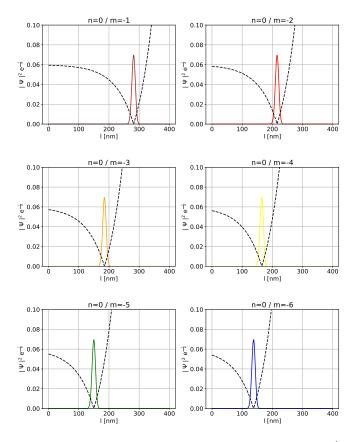


FIG. 4. Pseudosphere radial probability densities $|\Psi|^2 e^{-\frac{l}{a}}$ (solid lines, unnormalized) for selected zero-energy eigenstates of \hat{H}_{\perp} , Eq. (30), for different angular quantum numbers m. Probability density maxima of these zeroth Landau levels (n=0) are localized at the minima of the corresponding effective potentials $|V_{\perp}(l)|$ (dashed lines). Here: radius of curvature of pseudosphere $a=60\,\mathrm{nm}$ and perpendicular magnetic field $B=10\,\mathrm{T}$.

where $p_{\perp/\parallel}^{\uparrow}(\tilde{r})$ are so far unknown functions. Inserting the above ansatz into Eq. (53) gives the following differential equations for $p_{\perp/\parallel}^{\uparrow}(\tilde{r})$:

$$\tilde{r}^2 \frac{d^2 p_{\perp}^{\uparrow}}{d\tilde{r}^2} + \left[\left(1 - \frac{2}{\tilde{l}_B^2} \right) \tilde{r} - 2 \left(m + \frac{1}{2} \right) \right] \frac{d p_{\perp}^{\uparrow}}{d\tilde{r}}$$
(55)
$$+ E^2 p_{\perp}^{\uparrow} = 0 ,$$

$$\tilde{r}^2 \frac{d^2 p_{\parallel}^{\uparrow}}{d\tilde{r}^2} + \left[-\frac{1}{\tilde{l}_B^2} \tilde{r}^2 + \tilde{r} - 2\left(m + \frac{1}{2}\right) \right] \frac{dp_{\parallel}^{\uparrow}}{d\tilde{r}}$$

$$+ E^2 p_{\parallel}^{\uparrow} = 0$$
(56)

They will be solved for the different magnetic field configurations.

3. Higher Landau levels: perpendicular magnetic field

The solutions of Eq. (55) can be expressed in terms of the associated Laguerre polynomials of order $n \geq 0$ as shown in Appendix C. This constrains the square of the (dimensionless) energy E to satisfy

$$E_n^2 = n \left(\frac{2}{\tilde{l}_B^2} - n\right) = n \left(\frac{\Phi_\perp}{\Phi_0} - n\right). \tag{57}$$

Since only real eigenenergies are physical, we get an upper limit for the principal quantum number n defining the order of Laguerre polynomial. This limit is given by $\lfloor 2a^2/l_B^2 \rfloor = \lfloor \Phi_\perp/\Phi_0 \rfloor$, i.e. by a ratio of the squares of two characteristic length scales—the curvature radius a and the magnetic length l_B .

In view of the chiral symmetry, Eq. (35), the spectrum is symmetric with respect to zero, and hence the scaled eigenenergies E_n including the zero-mode read

$$E_n = \pm \sqrt{n\left(\frac{\Phi_{\perp}}{\Phi_0} - n\right)}, \quad n \in \left\{0, \dots, \left\lfloor \frac{\Phi_{\perp}}{\Phi_0} \right\rfloor \right\}. \quad (58)$$

The up components of the corresponding eigenfunctions read

$$\Psi_{\perp,n}^{\uparrow}(\tilde{r}) = \tilde{r}^n L_n^{(-2n+2/\tilde{l}_B^2)} \left(-\frac{2m+1}{\tilde{r}} \right) \Psi_{\perp,0}^{\uparrow}(\tilde{r}). \quad (59)$$

The down components are calculated by explicitly evaluating Eq. (45),

$$\Psi_{\perp,n}^{\downarrow}(\tilde{r}) = \frac{1}{E_n} \hat{L}_{\perp}^{+} \Psi_{\perp,n}^{\uparrow}(\tilde{r}), \qquad (60)$$

yielding again solutions in terms of an associated Laguerre polynomial:

$$\Psi_{\perp,n}^{\downarrow}(\tilde{r}) = i \frac{1}{E_n} \left(-n + \frac{2}{\tilde{l}_B^2} \right) \tilde{r}^n$$

$$\times L_{n-1}^{(-2n+2/\tilde{l}_B^2)} \left(-\frac{2m+1}{\tilde{r}} \right) \Psi_{\perp,0}^{\uparrow}(\tilde{r}).$$
(61)

Solutions for negative magnetic fields follow from the $\sigma_y\mathscr{C}$ symmetry, Eq. (36). The latter exchanges up and down components of the wavefunctions and reverses the signs of the angular momentum quantum number m; otherwise, the spectrum remains the same.

To summarize, independent of the sign of the magnetic field, the eigenenergies are degenerate in m and are given by (in units of $\mathcal{E} = \hbar v_F/a$)

$$E_n = \pm \sqrt{n \left(\frac{|\Phi_{\perp}|}{\Phi_0} - n \right)}, \quad n \in \left\{ 0, \dots, \left\lfloor \frac{|\Phi_{\perp}|}{\Phi_0} \right\rfloor \right\}.$$
(62)

However, as mentioned above, physically allowed solutions only exist for $|m| \leq |m_{crit,\perp}(B)|$.

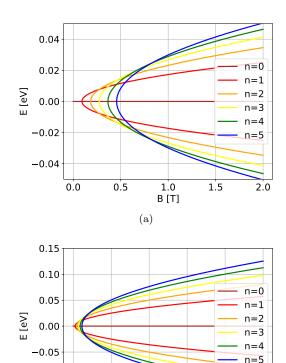


FIG. 5. Lowest Dirac Landau levels E_n , Eq. (62), for the pseudosphere in a perpendicular magnetic field \vec{B}_{\perp} as functions of B. Panel (a) shows a detailed view for small magnetic fields, while panel (b) offers a broader overview for larger fields. The pseudosphere radius of curvature $a=60\,\mathrm{nm}$.

(b)

B [T]

8

10

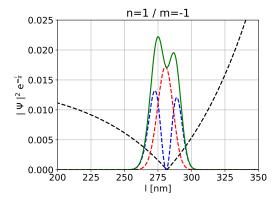
-0.10

-0.15

Fig. 5 shows the lowest eigenenergies as a functions of the magnetic field. Correspondingly, Fig. 6 displays the radial probability densities for the wavefunctions of the first two excited states with m=-1. It can be seen that for positive B the up component of the n-th excited state has n zeros—associate Laguerre polynomial of the n-th degree—and the down component (n-1) zeros—associate Laguerre polynomial of the (n-1)-th degree. Furthermore, it seems that the wavefunctions become spatially more broadened upon rising the principal quantum number n. We take this observation as numerical evidence without any rigorous proofs for a generic

4. Higher Landau levels: coaxial magnetic field

The experimentally more relevant case of a homogeneous coaxial magnetic field is more complicated since the B-field appears inhomogeneous on the surface of the pseudosphere and its perpendicular component is arclength-dependent. As shown in Appendix \mathbb{C} , the



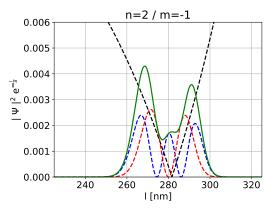


FIG. 6. Pseudosphere radial probability densities $|\Psi|^2 e^{-\frac{l}{a}}$ for the up component (blue, dashed lines), down component (red, dashed lines) and the whole spinor (green, solid lines) for the first and second excited states of \hat{H}_{\perp} , Eq. (30), with m=-1. Maxima of the wavefunction probabilities (not-normalized) are localized near the minima of the corresponding effective potentials $|V_{\perp}(l)|$ (black, dashed lines). The radius of curvature $a=60\,\mathrm{nm}$ and the strength of the perpendicular magnetic field $B=10\,\mathrm{T}$.

eigenfunctions of Eq. (56) can be expressed by means of the double confluent Heun functions. Since it is a bit cumbersome to work with these functions, we do not give explicit expressions for the underlying wavefunctions here; rather, we discuss the spectrum. Using asymptotic methods developed for the double confluent Heun equation on the complex-plane, see [78, 79] and references therein, the squares of eigenenergies E_n for $\tilde{l}_B^2 > 0$ and thus m < 0, recall the discussion in Sec. IV A – can be estimated by successive approximations, to the second order they are given by the following expression (normalized to $\mathcal{E} = \hbar v_F/a$)

$$E_{n,m}^2 \approx 2n\sqrt{-\frac{2m+1}{\tilde{l}_B^2}} + \frac{1}{2}n^2 - 1.$$
 (63)

The approximation works for high values of the argument under the square root, i.e., large |m| and large magnitude of B. Evidently, positive \tilde{l}_B^2 again require negative angular quantum numbers m. In accordance with

the chiral symmetry, Eq. (35), the eigenenergies are distributed symmetrically around zero energy. Expressing \tilde{l}_B^2 in terms of the perpendicular flux, Eq. (34), we get

$$E_{n,m} = \pm \sqrt{2n\sqrt{|(2m+1)|\frac{|\Phi_{\parallel}|}{\Phi_0}} + \frac{1}{2}n^2 - 1}$$
 (64)
$$\underset{n\gg 1}{\approx} \pm \sqrt{\frac{n}{2}\left(4\sqrt{|(2m+1)|\frac{|\Phi_{\parallel}|}{\Phi_0}} + n\right)}.$$

Here, it is worth to spell out several observations:

- (i) As can be seen, the approximate formula for $E_{n,m}$ does not include the zero energy eigenstates that are degenerate in m.
- (ii) Contrary to that and also to the case of an overall perpendicular field the higher Landau levels for \vec{B}_{\parallel} are no longer degenerate in m. This can be related to the fact that the eigenfunctions for different m are spatially centered near different l, the minima of $|V_{\parallel}|$, and the field strength component of \vec{B}_{\parallel} perpendicular to the surface Σ at l, i.e. $(\vec{n} \cdot \vec{B}_{\parallel})$, varies substantially with the arclength coordinate.
- (iii) Most notably, for strong magnetic fields, $E_{n,m} \propto |B|^{1/4}$. This peculiar parametric B-dependence should be experimentally observable. It characteristically differs from the usual linear scaling with |B| in the planar Schrödinger case, and also from the characteristic $\sqrt{|B|}$ -dependence for the corresponding Dirac case in constant perpendicular field.
- (iv) Moreover, due to the "+"-sign in front of the term $n^2/2$, and contrary to the case of \vec{B}_{\perp} , Eq. (58), n has no formal upper limit. However, when demanding that the states are localized within the surface, it becomes clear that the eigenenergies must not be larger than the confining effective potential, which sets a natural upper bound for n.

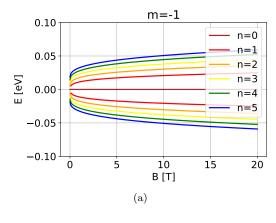
By comparing with numerical data in Sec. V we will show that the approximate expression for $E_{n,m}$, Eq. (64), is very accurate, for all m and $n \neq 0$.

C. Minding surface

In the following we sumarize the main spectral findings for the Minding surface depicted in Fig. 1.

1. Zeroth Landau level

Considering $\gamma \neq 0$ in the Eqs. (44) and (45) for the effective potentials $V_{\perp/\parallel}$ and the auxiliary operators $L_{\perp/\parallel}^{\pm}$ leads to slightly more involved radial Eq. (48). Introducing the auxiliary variable $\rho = \arcsin{(\tilde{r}/\tilde{\gamma})}$ the expressions



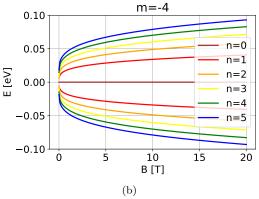


FIG. 7. Energies $E_{n,m}$ of the lowest Dirac Landau levels for a pseudosphere with a=60 nm in a coaxial magnetic field $\vec{B}_{\parallel}=(0,0,B)$, displayed as functions of B. Panels (a) and (b) show representative cases for angular quantum numbers m=-1 and m=-4, respectively. Results for finite n are based on the asymptotic expression, Eq. (64).

get simplified, and after some math the full spinorial solutions $\left(\Psi^{\uparrow}(\tilde{r},\phi),\Psi^{\downarrow}(\tilde{r},\phi)\right)^{\top}$ for the zeroth Landau level on the Minding surface for perpendicular and coaxial fields read

$$\Psi_{0,\perp}^{+} = \tilde{r}^{-\frac{1}{2} - \frac{1}{l_B^2}} e^{+\frac{1}{\tilde{\gamma}} \left(m + \frac{1}{2}\right) \arcsin\left(\frac{\tilde{\gamma}}{\tilde{r}}\right) + i\left(m + \frac{1}{2}\right)\phi} \begin{pmatrix} 1\\0 \end{pmatrix}, \tag{65}$$

$$\Psi_{0,\perp}^{-} = \tilde{r}^{-\frac{1}{2} + \frac{1}{\tilde{l}_B^2}} e^{-\frac{1}{\tilde{\gamma}} \left(m + \frac{1}{2}\right) \arcsin\left(\frac{\tilde{\gamma}}{\tilde{r}}\right) + i\left(m + \frac{1}{2}\right)\phi} \begin{pmatrix} 0\\1 \end{pmatrix}. \tag{66}$$

$$\Psi_{0,||}^{+} = \tilde{r}^{-\frac{1}{2}} e^{-\frac{\sqrt{\tilde{r}^{2} - \tilde{\gamma}^{2}}}{2\tilde{l}_{B}^{2}}} e^{+\frac{1}{\tilde{\gamma}} \left(m + \frac{1}{2}\right) \arcsin\left(\frac{\tilde{\gamma}}{\tilde{r}}\right) + i\left(m + \frac{1}{2}\right)\phi} \begin{pmatrix} 1\\0 \end{pmatrix}, \tag{67}$$

$$\Psi_{0,||}^{-} = \tilde{r}^{-\frac{1}{2}} e^{+\frac{\sqrt{\tilde{r}^2 - \tilde{\gamma}^2}}{2\tilde{l}_B^2}} e^{-\frac{1}{\tilde{\gamma}} \left(m + \frac{1}{2}\right) \arcsin\left(\frac{\tilde{\gamma}}{\tilde{r}}\right) + i\left(m + \frac{1}{2}\right)\phi} \begin{pmatrix} 0\\1 \end{pmatrix}. \tag{68}$$

Since Minding surfaces do not behave singularly and $r_{min} = \gamma$, both solutions are square-integrable. More-

over, in the limit $\gamma \to 0$ one indeed recovers the zeroth Landau level solutions for the pseudosphere given by Eqs. (49), (50), (51) and (52).

2. Higher Landau levels

Unfortunately, employing the same techniques for the Minding surface as for the pseudosphere does not yield any closed-form solution. Therefore, we resort to numerics. As shown and discussed below, we find strong numerical evidence that the energies of higher Landau levels for the Minding surface in \vec{B}_{\perp} are given by the same expression as Eq. (62) for the pseudosphere.

This is in accordance with the fact that the area of the embeddable part of the Minding surface with radius of curvature a does not depend on γ and becomes equal to the area of the pseudosphere possessing the same radius a and, hence, the total enclosed flux $|\Phi| = 2\pi a^2 |B|$.

V. NUMERICAL CALCULATIONS

Our analytical results for the pseudosphere in a coaxial field are only asymptotically exact, and we did not manage to solve the Dirac equation for the Minding surface in $\vec{B}_{\perp/\parallel}$ for non-zero energy states. We thus perform numerical calculations following the procedure laid out in Refs. [35, 36], adapted to the present context. For the sake of a self-contained presentation we recall its essentials.

A. Method and implementation

The continuous Dirac equation, Eq. (16), is discretized on a rectangular lattice and treated by means of the tight-binding package Kwant [80]. To deal with operators and fields in curvilinear coordinates, the tight-binding-generating hoppings are non-homogeneous and depend on entries of the metric tensor $g_{\mu\nu}(\phi,r)=\mathrm{diag}(r^2,\frac{a^2}{r^2-\gamma^2})$. One way to simplify this is to use the angle-arclength coordinates (ϕ,l) . Transforming from (ϕ,r) to (ϕ,l) simplifies the spatial-metric tensor from $g_{\mu\nu}(\phi,r)=\mathrm{diag}(r^2,\frac{a^2}{r^2-\gamma^2})$ to $g_{\mu\nu}(\phi,l)=\mathrm{diag}(r^2(l),1)$. Consequently, hoppings along the l direction on the underlying lattice do not depend on the current surface position, in contrast to the (ϕ,r) parameterization. Of course for both parameterizations, the hoppings along the ϕ coordinate are position-dependent.

Changing r to l, according to Eq. (26), transforms the Hamiltonians $\hat{H}_{\perp/\parallel}$, Eqs. (30) and (31), and also the underlying volume form

$$\omega_g = \sqrt{|\det g_{\text{space}}|} \, d\phi \wedge dl = r(l) \, d\phi \wedge dl \,.$$
 (69)

Denoting the transformed Hamiltonian as $\hat{H}(\phi,l)$, the differential operators with respect to l and ϕ entering it should be substituted by corresponding operators for finite differences on a discrete lattice. Using the symmetric difference scheme inherent to the flat Euclidean space leads to hermitian matrices in the conventional sense [72]. However, applying that prescription to $\hat{H}(\phi,l)$ would lead to a problem – the matrix of the discretized Hamiltonian $\hat{H}(\phi,l)$ would not be hermitian. Hence, the hermiticity with respect to the curved volume form $\omega_g = r(l) \, d\phi \wedge dl$ is required:

$$\langle \psi_1, \hat{H}\psi_2 \rangle = \int dl d\phi \ r(l) \ (\overline{\psi}_1)^{\mathcal{I}} \ (\hat{H}\psi_2)_{\mathcal{I}}$$
(70)
$$= \int dl d\phi \ r(l) \ \overline{(\hat{H}\psi_1)^{\mathcal{I}}} \ (\psi_2)_{\mathcal{I}}$$
(71)
$$= \langle \hat{H}\psi_1, \psi_2 \rangle .$$

We circumvent this issue by redefining the spinor fields and, accordingly, the Hamiltonian $\hat{H}(\phi, l)$ such that

$$\langle \psi_{1}, \hat{H}\psi_{2} \rangle = \int dl d\phi \ r(l) \ (\overline{\psi}_{1})^{\mathcal{I}} \ (\hat{H}\psi_{2})_{\mathcal{I}}$$

$$= \int dl d\phi \ \overline{(\sqrt{r(l)}\psi_{1})^{\mathcal{I}}} \ \left(\sqrt{r(l)}\hat{H}\frac{1}{\sqrt{r(l)}} \sqrt{r(l)}\psi_{2}\right)_{\mathcal{I}}$$

$$= \int dl d\phi \ (\overline{\widetilde{\psi}}_{1})^{\mathcal{I}} \ (\widetilde{H}\widetilde{\psi}_{2})_{\mathcal{I}}. \tag{72}$$

The appropriately rescaled Hamiltonian

$$\widetilde{H}(\phi, l) = \sqrt{r(l)} \, \widehat{H}(\phi, l) \, \frac{1}{\sqrt{r(l)}} \tag{73}$$

can now be routinely discretized by the standard symmetric difference scheme leading to a hermitian matrix in the conventional (flat space) fashion. In what follows we calculate the spectra of such rescaled Hamiltonians for the pseudosphere and Minding surfaces. We also add a Wilson mass term to the Hamiltonian \widetilde{H} in order to avoid fermion doubling, for details see [72]. Furthermore we use the hard-wall boundary conditions along the l direction (see discussion below) and anti-periodic boundary conditions along ϕ . For the sake of compactness we do not provide the explicit expressions of the rescaled Hamiltonians.

The embeddable part of the pseudosphere is non-compact. Thus we can apply the discretization only to a certain finite domain \mathscr{D} of Σ . For the latter we choose a part close to the rim (trumpet) of the pseudosphere, i.e. \mathscr{D} covers the part of radii descending from $r_{max}=a$ up to some end radius r_e , or in the arclength parameterization, the lengths from l=0 to $l_e=a\ln{(a/r_e)}$. So the discrete space is a 2D rectangular grid within $[0,2\pi]\times[0,l_e]$. We use N_W lattice points sampling the interval $[0,2\pi]$ and fix the number of lattice points in $[0,l_e]$ by introducing a lattice constant b, see Fig. 9. Using the symmetric difference approximation the discretization of

the derivatives along the ϕ and l directions is given by

$$\frac{d\tilde{\Psi}_{i,j}}{d\phi} \approx \frac{\tilde{\Psi}_{i+1,j} - \tilde{\Psi}_{i-1,j}}{2\frac{2\pi}{N_{\text{tot}}}},\tag{74}$$

$$\frac{d\tilde{\Psi}_{i,j}}{dl} \approx \frac{\tilde{\Psi}_{i,j+1} - \tilde{\Psi}_{i,j-1}}{2b},\tag{75}$$

where i and j numerate the lattice points in ϕ and l direction, respectively.

In order to compare analytical and numerical results we consider a pseudosphere and a Minding surface with the radius of curvature $a=60\,\mathrm{nm}$, and, correspondingly, with $\gamma=0$ and $\gamma=1\,\mathrm{nm}$ [81].

B. Pseudosphere: numerical vs. analytical results

For the numerics we use, besides $a=60\,\mathrm{nm}$, as minimum radius $r_e=0.25\,\mathrm{nm}$, corresponding to $l_e=328.8\,\mathrm{nm}$. This assumption is justified for wavefunctions decaying sufficiently fast towards the conically narrowing pseudosphere singularity at $r\to 0$ $(l\to\infty)$, as anticipated by our analytical ansatz, see Eq. (53).

Employing the hard-wall boundary conditions at $l=l_e$ and also at l=0 we expect certain deviations between the numerical and analytical spectra for small magnetic fields, as can be anticipated from the shapes and positions of the effective potentials $|V_{\perp/\parallel}(l)|$ shown in Figs. 2 and 3. For weak fields the potential minima approach the rim of the pseudosphere and, correspondingly, the hard-wall boundary conditions at l=0 will affect the numerically-computed eigendata, as opposed to analytical solutions. However, for sufficiently large fields, when the minima of $|V_{\perp/\parallel}|$ are pronounced and sufficiently far away from the rim, the numerical and analytical results should coincide and the finite-size effects stemming from the boundary-conditions would diminish.

We start with the case of the perpendicular magnetic field \vec{B}_{\perp} . In Fig. 8 we compare the numerically computed spectrum (as a function of field strength B) with the analytical results based on Eq. (62) For the chosen parameter range we find that for B>2 T the numerical eigenenergies (black symbols) agree very well with the corresponding analytical results (solid lines). As expected, upon increasing the magnetic field, the numerical eigenvalues rapidly converge to the analytical Landau levels E_n given by Eq. (62), in the order given by their principal quantum number n. For B<2T we see substantial deviations in the energy spectrum attributed to the effects of boundary-conditions.

Furthermore, Fig. 9 shows numerically computed radial probabilities for three eigenstates $(\Psi^{\uparrow}, \Psi^{\downarrow})^{\top}$ of the discretized Dirac Hamiltonian \widetilde{H} along with their analytical counterparts, given by Eqs. (59) and (61). Also these results corroborate the excellent matching between numerics and analytics at high(er) magnetic fields.

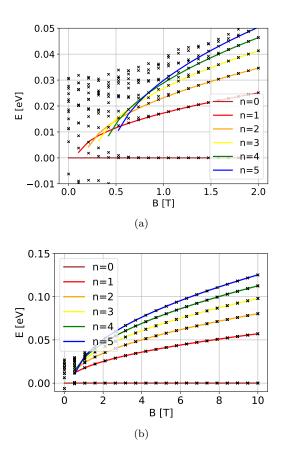


FIG. 8. Landau level spectrum of the pseudosphere in an overall perpendicular field \vec{B}_{\perp} of strength B. Low-lying eigenenergies $E_n(B)$ obtained analytically, Eq. (62), solid lines (color indicates Landau level quantum number n) are compared to numerical data (black symbols). Panel (a) shows a zoom into the low-field regime of the spectrum shown in panel (b) for larger B-scales. Numerical parameters used: radius of curvature $a=60\,\mathrm{nm}$, number of angular grid points $N_W=8$, and lattice spacing $b=0.5\,\mathrm{nm}$. Note that not all allowed higherlying eigenvalues are depicted and, for the sake of clarity, only zero and positive eigenvalues are displayed.

For the coaxial field \vec{B}_{\parallel} we employ the same discretization method. Here, however, the eigenenergies are not degenerate in the angular quantum number m, so fitting the numerical data with the help of the asymptotic eigenenergy formula, Eq. (64), would be quite messy and non-illustrative due to a substantial number of scattered points. For this reason we decided to only implement the radial equations for $(\Psi^{\uparrow}, \Psi^{\downarrow})^{\top}$, see Eqs. (46) and (47), for which the angular number m enters via the effective wedge-potential $|V_{\parallel}|$, consult Eq. (39) and Fig. 3. Fig. 10 shows the resulting numerical spectra for two different angular quantum numbers m as functions of the strength $B = ||\vec{B}_{\parallel}||$. It can be seen that despite the asymptotic eigenvalue formula, Eq. (64), the matching with the numerically obtained spectra becomes very good and improves for larger magnetic fields, higher angular and principal quantum numbers |m| and n, respectively. Fig. 11 shows the corresponding wavefunctions

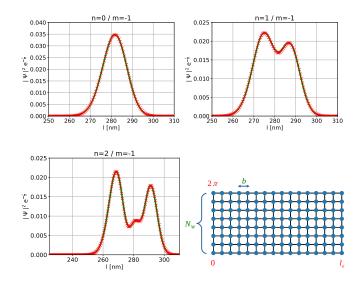
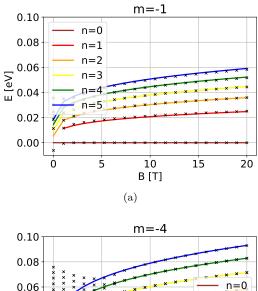


FIG. 9. Radial probability densities (unnormalized) for a pseudosphere in perpendicular magnetic field \vec{B}_{\perp} . Analytically obtained results, $|\Psi|^2 \, e^{-\frac{1}{a}}$, (solid, green lines) are compared to the corresponding numerical solutions (red crosses) for the three lowest Dirac Landau levels of \hat{H}_{\perp} with the angular quantum number m=-1. Numerical parameters: $N_W=20,\ b=0.5\,\mathrm{nm}$, pseudosphere radius $a=60\,\mathrm{nm}$, the field strength $B=10\,\mathrm{T.}$ – Schematic of the lattice used in the discretization of $\hat{H}_{\perp/\parallel}$ with N_W points along the angular direction, and length l_e and lattice constant b along the horizontal arclength direction.

for a few Landau levels in the coaxial field. As explained before, we only provide numerical solutions for the associated eigenfunctions, the analytical treatment being fairly complicated. Nevertheless, the probability maxima of the eigenstates are developing close to positions, where the effective wedge-potentials $|V_{\parallel}|$ become minimized, see Fig. 3.

C. Minding surface: numerical results

For the Minding surface we follow the same procedure as before; for the numerics we use $\gamma = 1 \,\mathrm{nm}$ and the same radius of curvature $a = 60 \,\mathrm{nm}$ as before. As the embeddable part of the Minding surface is finite, the discretization lattice samples the whole surface Σ . In case of the Minding surface, numerical diagonalization of the Dirac eigenvalue problem corresponding to the Hamiltonian \vec{H} with magnetic field \vec{B}_{\perp} yields the eigenenergies shown in Fig. 12. For high magnetic fields the Landau level spectrum agrees with the analytical results obtained for the pseudosphere in the perpendicular field, Eq. (62), as both carry the same magnetic flux assuming equal a and B. Hence, we have strong numerical evidences that Eq. (62) also applies to the Landau level spectrum for Minding surface Σ with $\gamma \ll a$ in the perpendicular magnetic field.



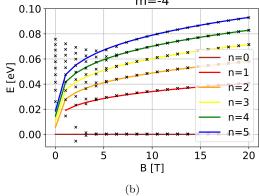


FIG. 10. Landau level spectrum of the pseudosphere in a coaxial field $\vec{B}_{\parallel}=(0,0,B)$. Lowest eigenenergies as functions of the magnetic field B obtained analytically, Eq. (64), (solid, colored lines labeled by principal quantum number) and numerically (black symbols) for m=-1, panel (a), and m=-4, panel (b). Here: the pseudosphere radius $a=60\,\mathrm{nm}$, and the arclength lattice constant $b=0.5\,\mathrm{nm}$, for the sake of brevity, only zero and positive eigenvalues are displayed.

Fig. 13 shows the numerical results for a coaxial field and angular quantum numbers m = -1 and m = -4, as in case of the pseudosphere. As for the perpendicular field, we fitted the data with the approximate dispersion $E_{n,m}$, Eq. (64), obtained for the pseudosphere. Matching is once more very good.

VI. CONCLUSION AND OUTLOOK

We analytically and numerically analyze (integer) quantum Hall states of massless Dirac fermions on prominent surfaces of revolution with constant negative Gaussian curvature – the pseudosphere and the Minding surface – subjected to perpendicular and coaxial magnetic fields. In special cases – zeroth Landau levels in $\vec{B}_{\perp/\parallel}$ and higher Landau levels for a pseudosphere in \vec{B}_{\perp} – we found explicit eigenfunctions. For all considered situations we provided numerical tight-binding calculations

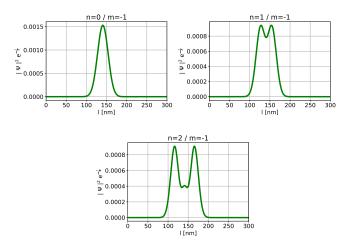


FIG. 11. Radial probability densities (unnormalized) obtained numerically for the three lowest (solid, green lines) pseudosphere Dirac Landau levels with the angular quantum number m=-1 in a coaxial magnetic field $\vec{B}_{\parallel}=(0,0,B)$. Numerical parameters: pseudosphere radius $a=60\,\mathrm{nm}$, arclength lattice constant $b=0.5\,\mathrm{nm}$, and magnetic field $B=20\,\mathrm{T}$.

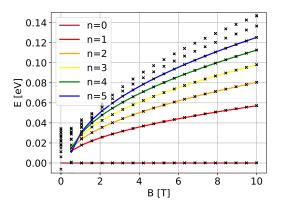


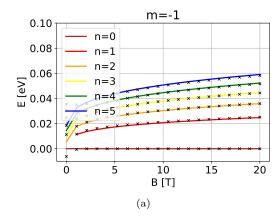
FIG. 12. Landau level spectrum of the Minding surface and pseudosphere in a perpendicular magnetic field \vec{B}_{\perp} . Lowest eigenenergies as functions of B are obtained numerically for the Minding surface (black symbols) and analytically for the pseudosphere (solid, colored lines), Eq. (62), both with the same radius of curvature a. Numerical parameters: $N_W=8$, the arclength lattice constant $b=0.5\,\mathrm{nm}$, radius of curvature $a=60\,\mathrm{nm}$ and $\gamma=1\,\mathrm{nm}$, for the sake of brevity, only zero and positive eigenvalues are shown explicitly.

that confirm our analytics with high accuracy.

In the following we compare our results from the spectral perspective and for different geometries, namely, for the flat and positively curved spaces with constant Gaussian curvature.

The Dirac Landau levels on the flat 2D plane—a surface with zero Gaussian curvature—in a homogeneous perpendicular field scales as

$$E_n = \pm C\sqrt{n|B|}, \text{ where } n = 0, 1, 2, \dots$$
 (76)



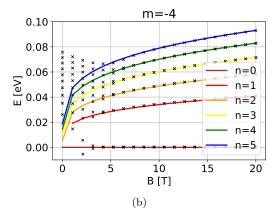


FIG. 13. Landau level spectrum of the Minding surface in a coaxial field $\vec{B}_{\parallel}=(0,0,B)$. Lowest eigenenergies of the pseudosphere as functions of the magnetic field B obtained analytically, Eq. (64), (solid, colored lines labeled by principal quantum number) lowest eigenenergies of the Minding surface obtained numerically (black symbols) for m=-1, panel (a), and m=-4, panel (b). Here: Minding surface radius of curvature $a=60\,\mathrm{nm}$ and $\gamma=1\,\mathrm{nm}$, the arclength lattice constant $b=0.5\,\mathrm{nm}$, for the sake of brevity, only zero and positive eigenvalues are displayed.

Counting both spin projections, each Landau level is degenerate by means of equal filling factor $|\Phi_{\perp}|/\Phi_0$, independent of the principal quantum number. Compared to other curved geometries, the zeroth Dirac Landau level in flat space is spin unpolarized.

Similarly, the energies of Dirac Landau levels on a 2D sphere with radius a in a perpendicular magnetic field (field of magnetic monopole) read

$$E_n = \pm \frac{\hbar v_F}{a} \sqrt{n \left(\frac{|\Phi_\perp|}{\Phi_0} + n\right)}, \text{ with } n = 0, 1, 2, \dots (77)$$

where $|\Phi_{\perp}|/\Phi_0$ again represents the number of flux quanta per a surface of the sphere [76, 77] with $|\Phi_{\perp}| = 4\pi a^2 |B|$. Although the Landau levels are specified by the number of flux quanta and the principal quantum number n, different Landau levels have different degeneracies, particularly, the n-th Landau level for $n \neq 0$ is $(2s_n + 1)$ degenerate, where $s_n = \lfloor |\Phi_{\perp}|/\Phi_0 \rfloor + n - 1/2$ and $\lfloor . \rfloor$

stands for the floor function. The zeroth Dirac Landau level is fully spin polarized, with spins parallel to $-\vec{B}_{\perp}$.

For the pseudosphere with curvature radius a we found both by analytical and numerical calculations that the Dirac Landau levels in a perpendicular field follow

$$E_n = \pm \frac{\hbar v_F}{a} \sqrt{n \left(\frac{|\Phi_\perp|}{\Phi_0} - n\right)} \quad . \tag{78}$$

Contrary to flat space and the positively curved sphere, the pseudosphere supports only a finite number of Landau levels with n running from 0 up to $|\Phi_{\perp}|/\Phi_{0}|$, with $|\Phi_{\perp}| = 2\pi a^2 |B|$. However, just as in flat space, the degeneracy of each Landau level n is given by a filling factor depending on the magnitude of B, but is independent of n, see Eq. (42). Note that the existence of a finite number of Landau levels can be guessed by inspection of the effective mass potential $|V_{\perp/\parallel}(l)|$, see Figs. 2 and 3: the potential does not diverge for $l \to 0$, yielding a well of finite depth. Moreover, we have very strong numerical evidences that the above formula also gives the correct Dirac Landau level spectrum for the Minding surfaces. Similarly to the case of a 2D sphere, the zeroth Dirac Landau level on the pseudosphere gets fully spin polarized, though now the spin points along the $+\vec{B}_{\perp}$ direction, i.e. the opposite compared to the 2D sphere.

Apart from the ranges of principal quantum numbers and different degeneracies of the Landau levels, the energy spectrum of massless Dirac electrons on a surface with constant Gaussian curvature κ in a perpendicular field \vec{B}_\perp reads in general

$$E_n = \pm \hbar v_F \sqrt{n(C_{\kappa}|B| + \kappa n)}. \tag{79}$$

It is worth to stress that the 2D sphere considered above is closed and compact, as opposed to the 2D plane and pseudosphere, therefore considering only its upper hemisphere C_{κ} becomes κ independent and equals $2\pi/\Phi_0$.

For the case of the perpendicular magnetic field \vec{B}_{\perp} we highlight the possibility to spectrally split the radial part of the Hamiltonian, Eq. (41), in terms of \hat{L}^{\pm} operators. This can be interpreted and approached from the point of view of 1D supersymmetric quantum mechanics [82, 83]. However, we did not follow this approach in our work, but as expected [83], the zeroth Landau level physics is connected to the index theorem of the corresponding Dirac operator.

A further central aspect of our work consists of the study of the coaxial magnetic field configuration, which to the best of our knowledge has hitherto not been considered. It is however particularly relevant since its experimental realization seems not too far, considering current capabilities in shaping TI nanowires [32–34], which host massless Dirac electrons on their (curved) surfaces. In such a configuration we did not succeed in finding exact analytical solutions to the problem, but via asymptotic methods we could determine approximate eigen-

spectra. Our results were fully confirmed by numerical tight-binding calculations.

The spectral structure of the pseudosphere and Minding surface in a coaxial field can be probed, e.g., optically or via transport. Moreover, it could have interesting implications for orbital magnetic phenomena. A general question is indeed how the magnetic properties of quantum systems in curved space may be affected by the local (Gaussian) curvature, and whether for example signatures of fractional angular momentum quantization, as proposed in Ref. [84], can be identified.

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Appendix A: Spin connection for pedestrians – adiabatic parallel transport on surfaces of revolution

2D surfaces of revolution Σ embedded in 3D Euclidean space parameterized by (ϕ,r) were discussed—including the tangent vectors $\vec{\partial}_{\phi}$, $\vec{\partial}_{r}$ and the induced metric tensor $g_{\mu\nu}$ —in section II. It is this Euclidean ambient space, and the fact that the spinor representations in 2D and 3D spaces are both two-dimensional, that allows one to define a spin connection in a very natural manner. However, the formal definition as provided by Eq. (3) is intrinsic, i.e. it is valid for any manifold and is independent of any embedding.

To describe spinor fields on Σ we need to specify a quantization axis, that we define point-wisely, i.e. at any point (ϕ,r) on Σ the local quantization axis is given by the unit vector \vec{n} , that corresponds to the outer normal $\vec{n} \propto \vec{\partial}_{\phi} \times \vec{\partial}_{r}$ of the surface Σ at that point. Here \times is the standard vector product in the 3D ambient Euclidean space. Evaluating \vec{n} , using expressions given by Eqs. (8), one gets:

$$\vec{n} = \frac{1}{\sqrt{1 + \left(\frac{dz}{dr}\right)^2}} \left(\frac{dz}{dr}\cos\phi, \frac{dz}{dr}\sin\phi, 1\right).$$
 (A1)

Since $1 + (dz/dr)^2 \ge 1$ one can define an auxiliary angle ϑ that depends on r and dz/dr, such that

$$\cos \vartheta = \frac{1}{\sqrt{1 + \left(\frac{dz}{dr}\right)^2}} = \hat{\omega}_{\phi}^{12} \iff \sin \vartheta = \frac{\frac{dz}{dr}}{\sqrt{1 + \left(\frac{dz}{dr}\right)^2}},$$
(A2)

or $\tan \vartheta = dz/dr$. Doing so one can rewrite the normal vector in a conventional form:

$$\vec{n} = (\cos\phi\sin\theta, \sin\phi\sin\theta, \cos\theta). \tag{A3}$$

Hence, specifying \vec{n} at a given point (ϕ, r) means to define point-wisely a spin basis that is framed by spin up and spin down states, $|+\rangle_{(\phi,r)}$ and $|-\rangle_{(\phi,r)}$ —which are chosen as the eigenstates of $\vec{n} \cdot \vec{\sigma}$. Using $\cos \vartheta$ and $\sin \vartheta$ that depend on r and dz/dr, and coordinate ϕ we can write them in a conventional and very compact way:

$$|+\rangle_{(\phi,r)} = \begin{pmatrix} +e^{-i\phi/2}\cos\vartheta/2\\ e^{+i\phi/2}\sin\vartheta/2 \end{pmatrix},$$
 (A4)

$$|-\rangle_{(\phi,r)} = \begin{pmatrix} -e^{-i\phi/2} \sin \vartheta/2 \\ e^{+i\phi/2} \cos \vartheta/2 \end{pmatrix}. \tag{A5}$$

This is our particular choice (gauge) of the spin bundle trivialization over the surface Σ . Someone else can, however, impose a different gauge, i.e. instead of $|\pm\rangle_{(\phi,r)}$, use $e^{if_{\pm}(\phi,r)}|\pm\rangle_{(\phi,r)}$ with some locally defined phasefunctions $f_{\pm}(\phi,r)$ (no global definition over the whole Σ is required). Moreover, observe that we are picking the "anti-periodic gauge" i.e.

$$|\pm\rangle_{(\phi+2\pi,r)} = -|\pm\rangle_{(\phi,r)}. \tag{A6}$$

Having a spinor-field Ψ on Σ , therefore, means to define ϕ -anti-periodic component fields $\Psi_{\uparrow}(\phi, r)$ and $\Psi_{\downarrow}(\phi, r)$, such that the full spinor

$$\Psi = \Psi_{\uparrow}(\phi, r) |+\rangle_{(\phi, r)} + \Psi_{\downarrow}(\phi, r) |-\rangle_{(\phi, r)}
\equiv \begin{pmatrix} \Psi_{\uparrow}(\phi, r) \\ \Psi_{\downarrow}(\phi, r) \end{pmatrix},$$
(A7)

is a global object unambiguously defined over the surface Σ . It is this anti-periodic gauge what imposes the ansatz, Eq. (37), taking the angular part of Ψ in the form $e^{i(m+1/2)\phi}$.

The parallel transport is a rule prescribing to what spinor will correspond the spinor $|\pm\rangle_{(\phi-d\phi,r-dr)}$ when moved from the given point $(\phi-d\phi,r-dr)$ to an infinitesimally close point (ϕ,r) . Let us denote that parallerly transported spinor at (ϕ,r) as $PT|\pm\rangle_{(\phi,r)}$, and let us decompose it with respect to the corresponding $|\pm\rangle_{(\phi,r)}$ states residing at (ϕ,r) , i.e.

$$PT|\pm\rangle_{(\phi,r)} = \mathcal{A}_{\pm}(\phi,r)|+\rangle_{(\phi,r)} + \mathcal{B}_{\pm}(\phi,r)|-\rangle_{(\phi,r)}.$$
(A8)

So technically, the parallel transport is given by a prescription for the position dependent coefficients \mathcal{A}_{\pm} and \mathcal{B}_{\pm} .

From the point of view of 2D surface, the spinors $|\pm\rangle_{(\phi-d\phi,r-dr)}$ and $|\pm\rangle_{(\phi,r)}$, which are residing at different (although nearby) points do not know about each other, so something like their sum, or inner product do not have a priori any sense. However, from the point of view of the ambient 3D Euclidean space and its flat spin structure one can move $|\pm\rangle_{(\phi-d\phi,r-dr)}$ and $|\pm\rangle_{(\phi,r)}$ to the origin of \mathbf{R}^3 and perform their scalar product there etc. The rule we prescribe for $\mathrm{PT}|\pm\rangle$ —i.e. rules for \mathcal{A}_{\pm} and \mathcal{B}_{+} —relies on that ambient space, namely, we define

$$\mathcal{A}_{+} = \langle \langle +_{(\phi,r)} | +_{(\phi-d\phi,r-dr)} \rangle \rangle \simeq 1 - i \frac{d\phi}{2} \cos \vartheta , \quad (A9)$$

$$\mathcal{A}_{-} = \langle \langle +_{(\phi,r)} | -_{(\phi-d\phi,r-dr)} \rangle \rangle \simeq 0, \tag{A10}$$

$$\mathcal{B}_{+} = \langle \langle -_{(\phi,r)} | +_{(\phi-d\phi,r-dr)} \rangle \rangle \simeq 0, \tag{A11}$$

$$\mathcal{B}_{-} = \langle \langle -_{(\phi,r)} | -_{(\phi-d\phi,r-dr)} \rangle \rangle \simeq 1 + i \frac{d\phi}{2} \cos \theta \,, \quad (A12)$$

where the meaning of \simeq on the right-hand sides means an expansion up to the first order in $d\phi$ and $d\vartheta \propto dr$. More importantly, the scalar product $\langle\langle \cdot | \cdot \rangle\rangle$ entering the above definition is understood in a sense of the standard two-component \mathbb{C}^2 Pauli spinors in the ambient Euclidean space \mathbb{R}^3 , and not in a sense of Eq. (4). Since the parallel transport preserves spin projections on infinitesimal distances, $\mathcal{A}_- = 0 = \mathcal{B}_+$, and preserves norms, $|\mathcal{A}_+| = 1 = |\mathcal{B}_-|$, both up to the first order in $d\phi$ and dr, it is called the *spin-adiabatic parallel-transport* or *spin connection*, but in an essence it is the Berry connection.

Covariant derivatives, or more precisely, their spinor part components—see Eqs. (13) and (14)—are defined as:

$$\nabla_{r}|\pm\rangle_{(\phi,r)} := \lim_{dr \to 0} \frac{|\pm\rangle_{(\phi,r)} - \operatorname{PT}|\pm\rangle_{(\phi,r)}}{dr}$$

$$= 0, \qquad (A13)$$

$$\nabla_{\phi}|\pm\rangle_{(\phi,r)} := \lim_{d\phi \to 0} \frac{|\pm\rangle_{(\phi,r)} - \operatorname{PT}|\pm\rangle_{(\phi,r)}}{d\phi}$$

$$= \pm \frac{i}{2}\cos\vartheta |\pm\rangle_{(\phi,r)}$$

$$= \frac{i}{2}\hat{\omega}_{\phi}^{12}\sigma_{z} |\pm\rangle_{(\phi,r)}, \qquad (A14)$$

what is exactly the spin connection we have encountered in section II. So ∇_r and ∇_ϕ "quantify" how much the spin-bundle trivialization, i.e. states $|\pm\rangle_{(\phi,r)}$, changes when transported infinitesimally between neighbouring points on Σ .

Appendix B: Christoffel symbols

An elegant way to obtain the Christoffel symbols for a given metric and curved coordinates is to use the Lagrange function—kinetic energy of a unit mass particle on the curved surface [65],

$$\mathcal{L}(\phi, \dot{\phi}, r, \dot{r}) = \frac{1}{2} (\dot{\phi}, \dot{r}) \begin{pmatrix} r^2 & 0 \\ 0 & 1 + \left(\frac{dz(r)}{dr}\right)^2 \end{pmatrix} \begin{pmatrix} \dot{\phi} \\ \dot{r} \end{pmatrix}, \quad (B1)$$

where the matrix represents the spatial part of the spacetime metric tensor $g_{\mu\nu}={\rm diag}(-v_F^2,r^2,1+\left(dz(r)/dr\right)^2)$ and the dot denotes the derivative with respect to a curve parameter, say time, on which both coordinates, ϕ and r depend upon. Writing down the corresponding Euler-Lagrange equation for ϕ one recovers the Christoffel symbols Γ_{-}^{ϕ} [65],

$$\frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \dot{\phi}} \right) - \frac{\partial \mathcal{L}}{\partial \phi} = 2r\dot{r}\dot{\phi} + r^2\ddot{\phi} = 0.$$
 (B2)

This is equivalent to

$$2\frac{1}{r}\dot{r}\dot{\phi} + \ddot{\phi} \stackrel{!}{=} \Gamma^{\phi}_{r\phi}\dot{r}\dot{\phi} + \Gamma^{\phi}_{\phi r}\dot{\phi}\dot{r} + \ddot{\phi} = 0, \qquad (B3)$$

out of which the non-zero Christoffel symbols read

$$\Gamma^{\phi}_{\phi r} = \Gamma^{\phi}_{r\phi} = \frac{1}{r} \,. \tag{B4}$$

Repeating the same for the r coordinate, one recovers the corresponding non-zero Γ^r Christoffel symbols:

$$\Gamma_{rr}^{r} = \frac{\frac{dz}{dr}\frac{d^{2}z}{dr^{2}}}{1 + \left(\frac{dz}{dr}\right)^{2}} \quad \Gamma_{\phi\phi}^{r} = -\frac{r}{1 + \left(\frac{dz}{dr}\right)^{2}}$$
(B5)

Appendix C: Solutions of the radial equations

The differential equation we need to solve for the $E \neq 0$ states in perpendicular magnetic field, see Eq. (55), is of the form

$$x^{2} \frac{d^{2}y(x)}{dx^{2}} + (\alpha x + \beta) \frac{dy(x)}{dx} + \gamma y(x) = 0,$$
 (C1)

and its coefficients α , β and γ follow from Eq. (55). This is a degenerate double-confluent Heun equation which can be simplified using the transformation [86]

$$\tilde{x} = \frac{\beta}{x}, \ y(\tilde{x}) = \tilde{x}^{\Delta_{\pm}} \tilde{y}(\tilde{x}), \ \Delta_{\pm} = \frac{\alpha - 1 \pm \sqrt{(\alpha - 1)^2 - 4\gamma}}{2}.$$
 (C2)

This leads to a confluent hypergeometric equation

$$\tilde{x}\frac{d^2\tilde{y}(\tilde{x})}{d\tilde{x}^2} + \left[(2\Delta_{\pm} + 2 - \alpha) - \tilde{x} \right] \frac{d\tilde{y}(\tilde{x})}{d\tilde{x}} - \Delta_{\pm}\tilde{y}(\tilde{x}) = 0.$$
(C3)

For a given Δ , this equation has two linearly independent solutions, for instance Kummer's function M(a,b,z) and Tricomi's function U(a,b,z). Counting the possibility to choose Δ_+ or Δ_- in the above transformation, this would seemingly yield four solutions. Of course, not all of them

can be linearly independent since the original equation is only of second order. A basis of solutions may be given by

$$y_{-}(x) = x^{-\Delta_{-}} M\left(\Delta_{-}, 2\Delta_{-} + 2 - \alpha, \frac{\beta}{x}\right),$$
 (C4)

$$y_{+}(x) = x^{-\Delta_{+}} M\left(\Delta_{+}, 2\Delta_{+} + 2 - \alpha, \frac{\beta}{x}\right).$$
 (C5)

Both solutions become polynomials in case the first argument of M becomes a non-positive integer. Before we examine this further, we want to show that for our problem only polynomial solutions lead to physical, i.e. normalizable, wavefunctions. For this we need to inspect the asymptotic behavior of the whole wavefunction in the limit $\tilde{r} \to 0$ and hence in particular M(a,b,z) for $z \to \infty$ and $z \to -\infty$, assuming that the first parameter is not a non-positive integer. The first limit, $z \to \infty$, is given by

$$z \to \infty : M(a, b, z) \to \frac{e^z z^{a-b}}{\Gamma(a)}.$$
 (C6)

For $z \to -\infty$, we use Kummer's transformation

$$M(a, b, z) = e^{z} M(b - a, b, -z)$$
 (C7)

which gives us the relation

$$z \to -\infty$$
: $M(a, b, z) = e^{-z}M(b - a, b, -z) \to \frac{z^{-a}}{\Gamma(b - a)}$. (C8)

Recall that in our original ansatz for $\Psi_{\perp}^{\uparrow}(\tilde{r})$, Eq. (53), we used a function which in the present context has asymptotic behavior $e^{-\frac{\beta}{2x}}$. So assuming β is positive (we need the limit $z \to \infty$), the overall behavior is $e^{+\frac{\beta}{2x}}$, which clearly blows up when $x \to 0$. In case of a negative β (we need the limit $z \to -\infty$), the asymptotic behavior is given by $e^{-\frac{\beta}{2x}}$, which also blows up. This clearly shows that only polynomial solutions yield physical wavefunctions. The condition that at least one of the two basis solutions behaves polynomially is given by

$$\Delta_{+} \stackrel{!}{=} -n, n \in \{0, 1, 2, \dots\},$$
 (C9)

what gives us the quantization condition for the eigenenergies. For given α and β , it turns out that this equation can only be solved for Δ_{-} . The solution reads

$$\gamma = -n^2 - n\left(\alpha - 1\right). \tag{C10}$$

Overall, our solution for the radial differential equation reads

$$y(x) = x^n M\left(-n, -2n + 2 - \alpha, \frac{\beta}{x}\right).$$
 (C11)

This can be simplified even further employing the associated Laguerre polynomials $L_n^{(a)}(x)$

$$y(x) = x^n L_n^{(-2n+1-\alpha)} \left(\frac{\beta}{x}\right),$$
 (C12)

where we have absorbed some constants into the overall normalization (not shown).

The differential equation needed for the coaxial magnetic field, Eq. (56), can be brought into the form

$$x^{2} \frac{d^{2}y(x)}{dx^{2}} + (\alpha x^{2} + x + \beta) \frac{dy(x)}{dx} + \gamma y(x) = 0, \quad (C13)$$

with the coefficients α , β and γ stemming from Eq. (56). This is a double-confluent Heun equation. Contrary to the previous case, it is not degenerate (for $\alpha \neq 0$, $\beta \neq 0$) and hence a transformation into the much simpler confluent hypergeometric differential equation is not possible. However, further insight can be gained when performing the rescaling

$$\tilde{x} = -\alpha x$$
, (C14)

which yields

$$\tilde{x}^2 \frac{d^2 y(\tilde{x})}{d\tilde{x}^2} + (-\tilde{x}^2 + \tilde{x} - \alpha \beta) \frac{dy(\tilde{x})}{d\tilde{x}} + \gamma y(\tilde{x}) = 0. \quad (C15)$$

Interestingly, the coefficients α and β do not enter independently into the equation. As in our original problem γ plays a role of the energy eigenvalue, the above equation then asserts that γ depends only on the product $\alpha\beta$ and not on their separate values. The solutions of the differential equation can be expressed in terms of the double-confluent Heun function. Since it is quite involved to work with them, we relax an attempt to provide the wavefunctions in analytic form and rather focus on the eigenenergies. To our best knowledge, there is currently no exact formula determining the spectrum of this equation. However, there are known asymptotic methods and approximations [79] which allow to get the eigenvalues expanded in reciprocal powers of $\sqrt{-\alpha\beta}$, in particular up to second order,

$$\gamma_n = 2n\sqrt{-\alpha\beta} + \frac{1}{2}n^2 - 1. \tag{C16}$$

Using this formula and expressing α and β in terms of the original parameters, one recovers the asymptotic eigenvalue results as given by Eq. (64).

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